QUCS12 December 14, 2012

Strangeness Nuclear Physics Present status and future prospect at J-PARC



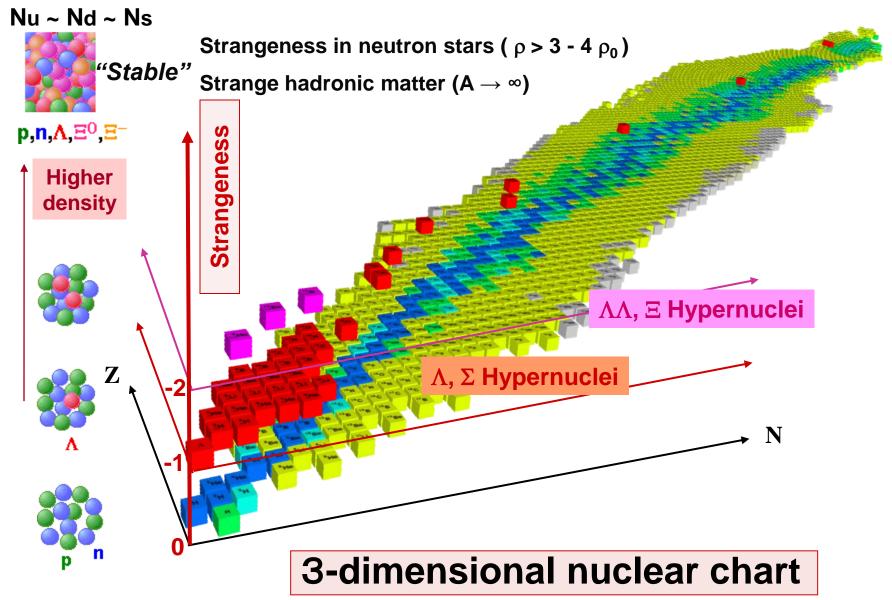
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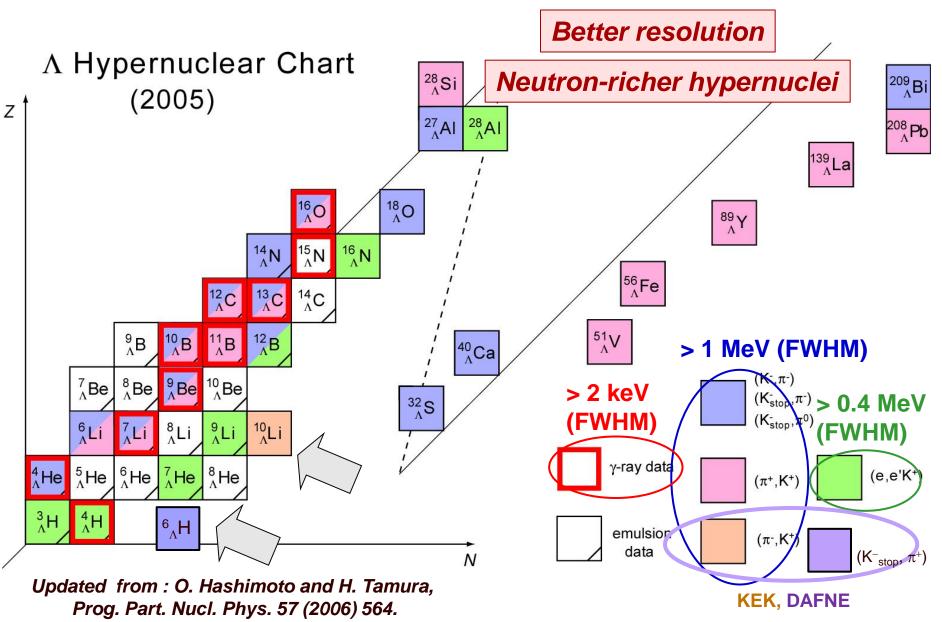
1. Introduction

World of matter made of u, d, s quarks



by M. Kaneta inspired by HYP06 conference poster

<u>Λ hypernuclear chart</u>



Motivations of strangeness nuclear physics

Extension of 3D nuclear chart

BB interactions

Unified understanding of BB forces by u,d ->u, d, s particularly short-range forces by quark pictures Test lattice QCD calculations

Impurity effect

in nuclear structure

Changes of size, deformation, clustering, collective motions, appearing new symmetry, Modifications of baryon properties

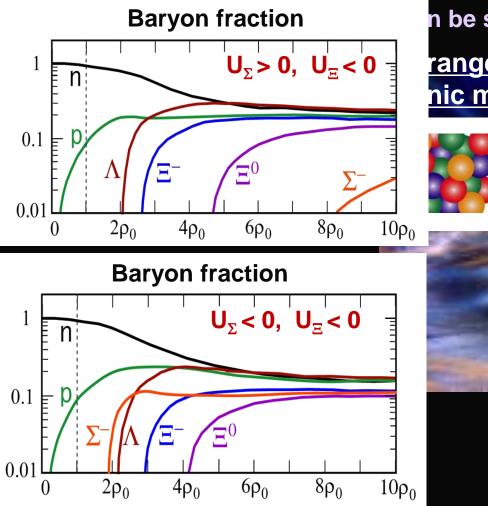
$\underbrace{\text{in nuclei}}_{\mu_{\Lambda} \text{ in a nucleus,}} \quad \stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}{\stackrel{\mu_{\Lambda}}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}{\stackrel{\mu_{\Lambda}}}}}}}}}$

Levels of heavy Λ hypernuclei,

Clues to understand hadrons and nuclei from quarks Cold and dense nuclear matter with strangeness



Nuclear matter in neutron stars



n be supported? **Neurtron matter** range nic matter **Quark matter**

We still need $\Xi N, \Sigma N, \Lambda \Lambda, K^{bar}N$ forces, ΛN odd-state force, NNN and YNN force, ... A new form of matter actually existing in the universe

Overview of

Strangeness Nuclear Physics Experiments

- K⁻ or π^{+/-} beams Intensity increasing.
 KEK-PS, BNL-AGS => J-PARC SNP programs just being started.
 Λ and Σ hypernuclear spectroscopy, γ-spectroscopy of Λ hypernuclei
 ΛΛ and Ξ hypernuclei,
 ΣN /ΛN scattering
 weak decays of Λ hypernuclei
 K⁻ nuclei, K⁻ atoms
- e⁻ beam

Jefferson Lab (Hall A, Hall C), MAMI-C

high-res. A hypernuclear spectroscopy, weak decays of A hypernuclei

K⁻ from φ

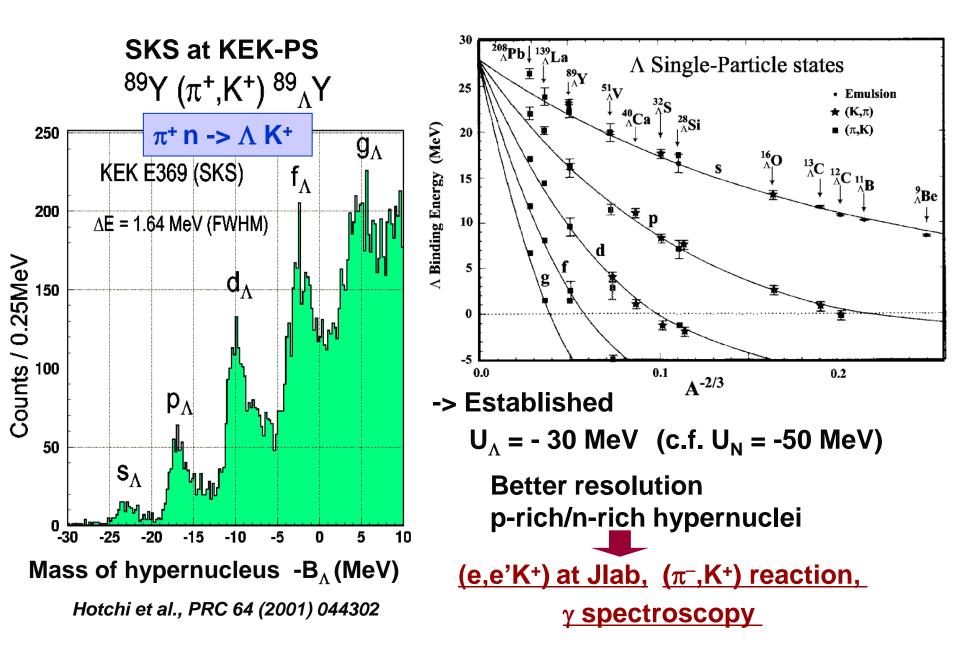
DAΦNE (FINUDA, SIDDHARTA, AMADEUS) spectroscopy of Λ hypernuclei, weak decays of Λ hypernuclei K⁻ nuclei, K⁻ atoms

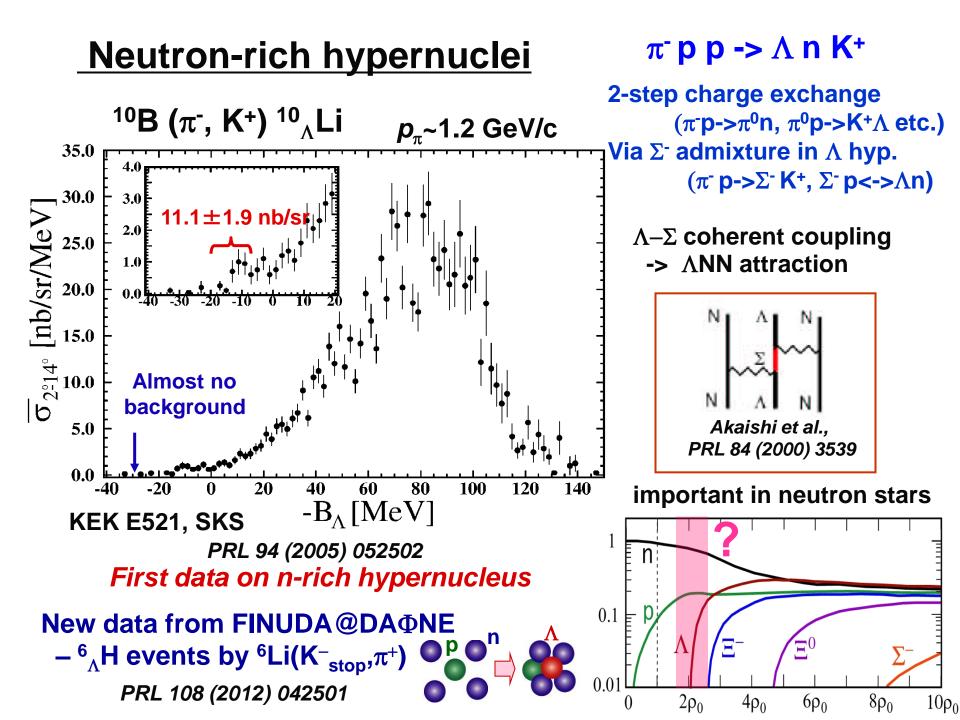
Heavy ion beams
 GSI (HypHI) p-rich/n-rich hypernuclei, lifetimes, multi-strange nuclei

2. A hypernuclei

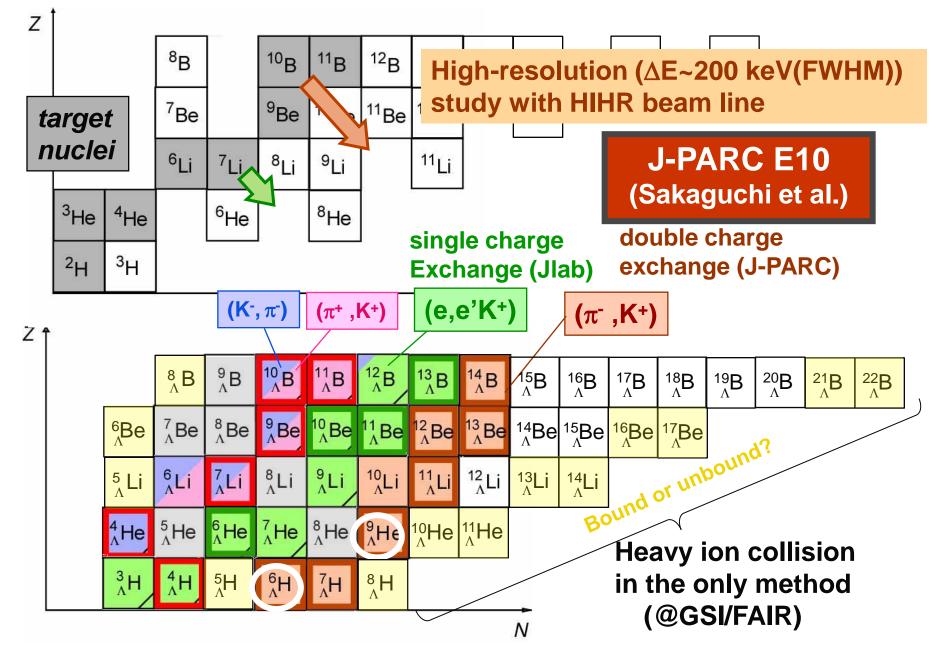
2.1 (π, K⁺) spectroscopy and n-rich Λ hypernuclei

Previous (π^+ ,K⁺) data and ΛN interaction



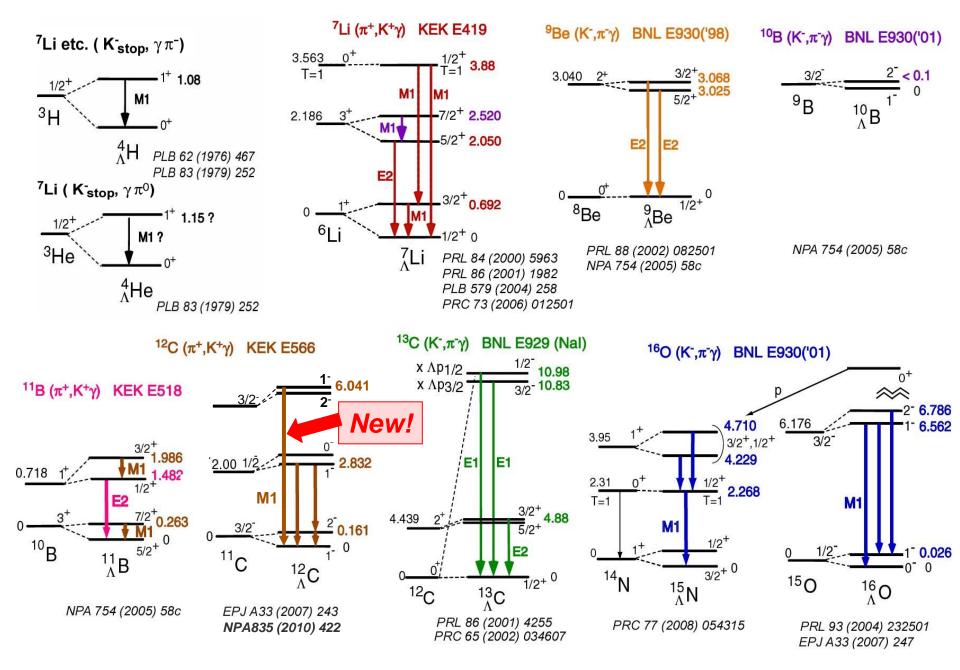


How to extend S=-1 nuclear chart?

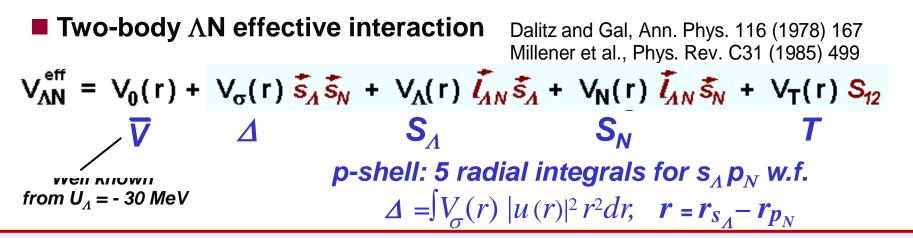


2.2 γ spectroscopy

Hypernuclear γ-ray data (2012)



<u>AN spin-dependent interactions</u>



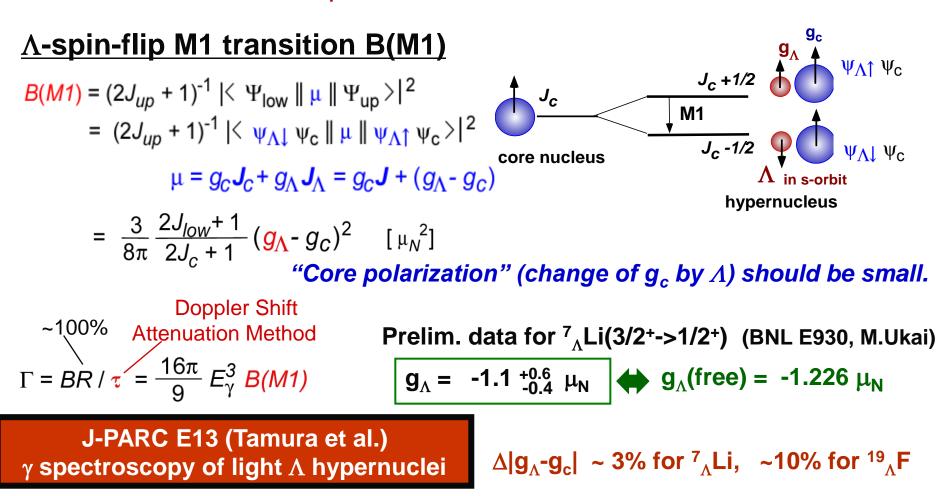
 γ -ray data => Δ = 0.33 (0.43 for A=7), S_A = -0.01, S_N = -0.4, T = 0.03 [MeV] Small spin-dependent forces have been established.

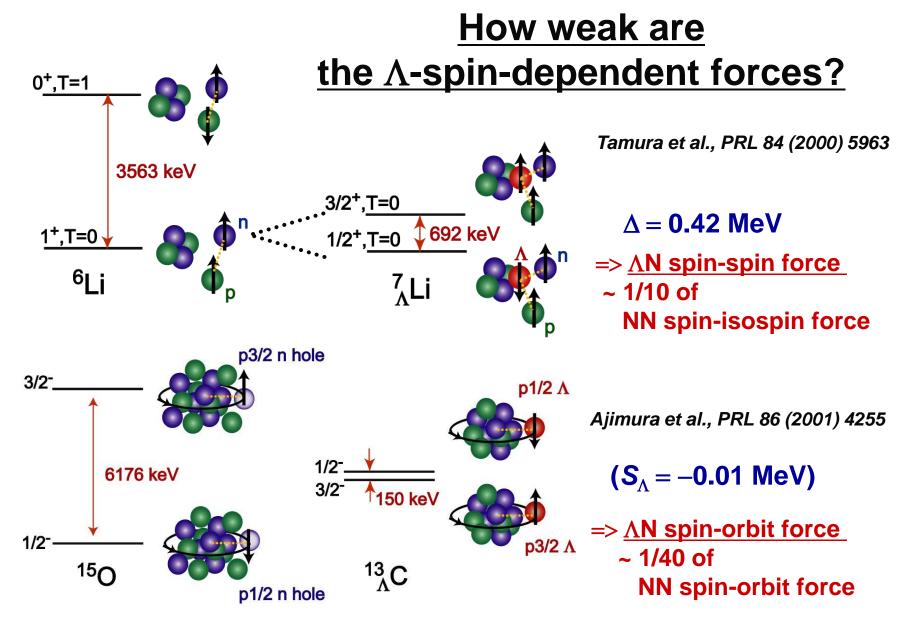
Feedback to		Δ	S _A	S _N	Τ	(MeV)
BB interaction	ND	-0.048	-0.131	-0.264	0.018	
models thru	NF	0.072	-0.175	-0.266	0.033	LS force:
G-matrix calc.	NSC89	1.052	-0.173	-0.292	0.036	All Nijmegen
(Millener)	NSC97f	0.421	-0.149	-0.238	0.055	models fail.
	ESC04a	0.381	-0.108	-0.236	0.013	Quark model
Nijmegen models	ESC08a	0.146	-0.074	-0.241	0.055	looks OK.
	("Quark model"		0.0	-0.4)		
	Exp.	0.4	-0.01	-0.4	0.03	

g-factor of Λ in a nucleus

Changes by partial restoration of chiral symmetry? $\mu_{q} = \frac{e \hbar}{2m_{q}c} \qquad m_{q}: \text{Const. quark mass}$

Reduction of m_q in nuclear matter -> enhancement of μ ??





But the core polarization effect should be theoretically estimated. + Meson exchange effect + Σ - Λ mixing effects should be estimated.

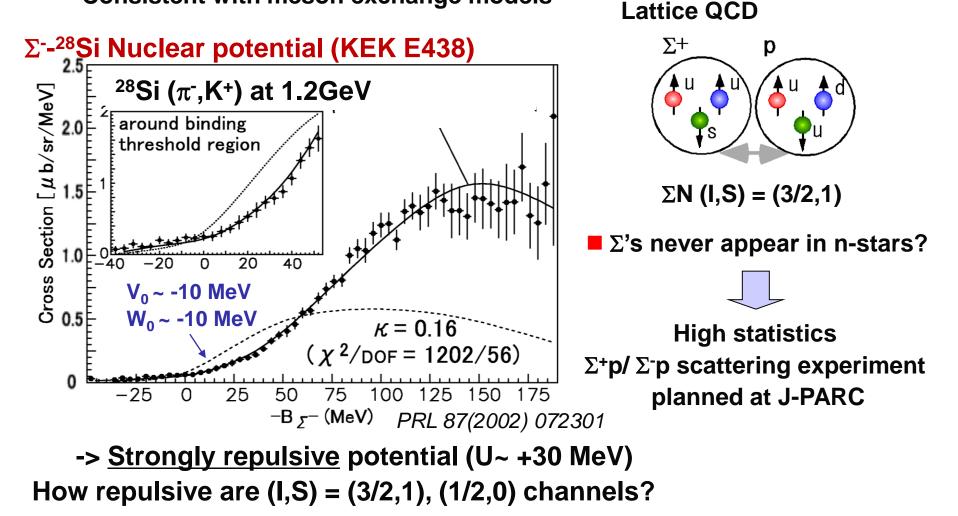
3. Σ-nuclear systems

What we know about Σ –N force

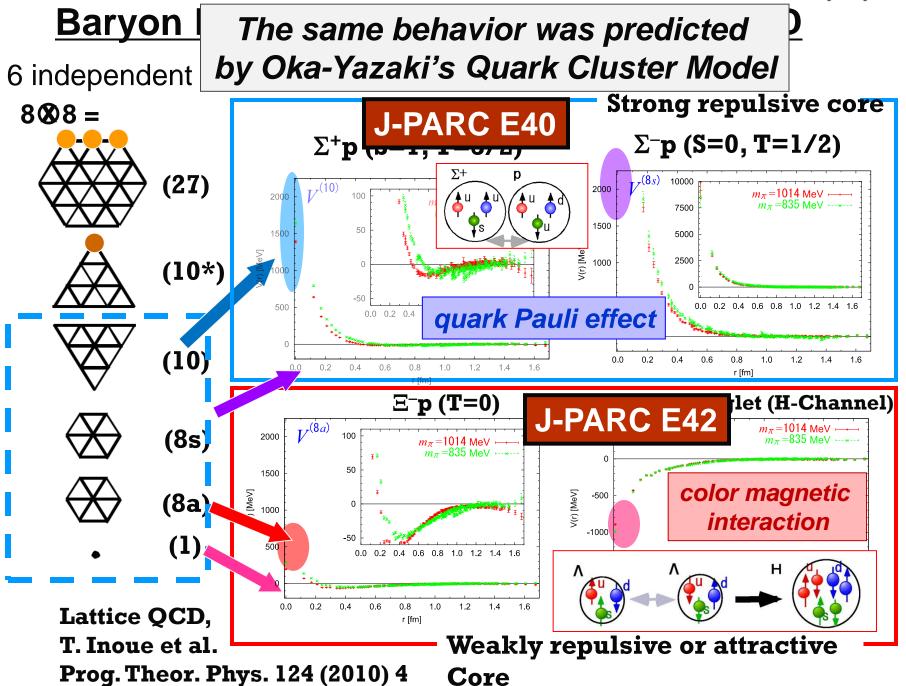
Strong repulsion comes from

Quark Cluster Model

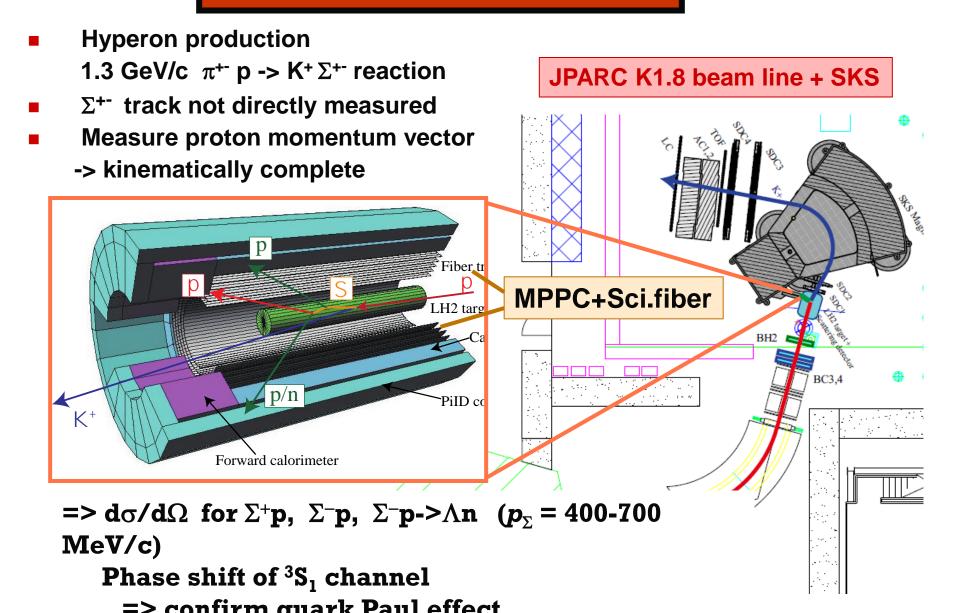
Large spin-isospin dependence in ΣN force ${}^{4}_{\Sigma}$ He suggests (I,S) = (3/2,0), (1/2,1) attractive Pauli effect between quarks? (3/2,1), (1/2,0) repulsive -- Consistent with meson exchange models



Slide by Koji Miwa



J-PAR E40 (Miwa et al.) Σp Scattering Experiment

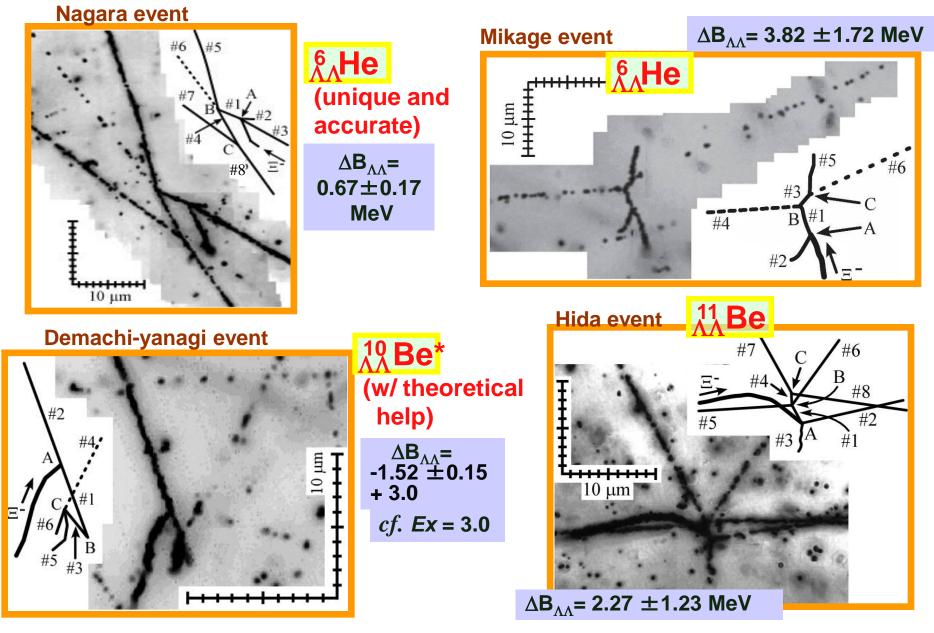


4. Double strange nuclear systems

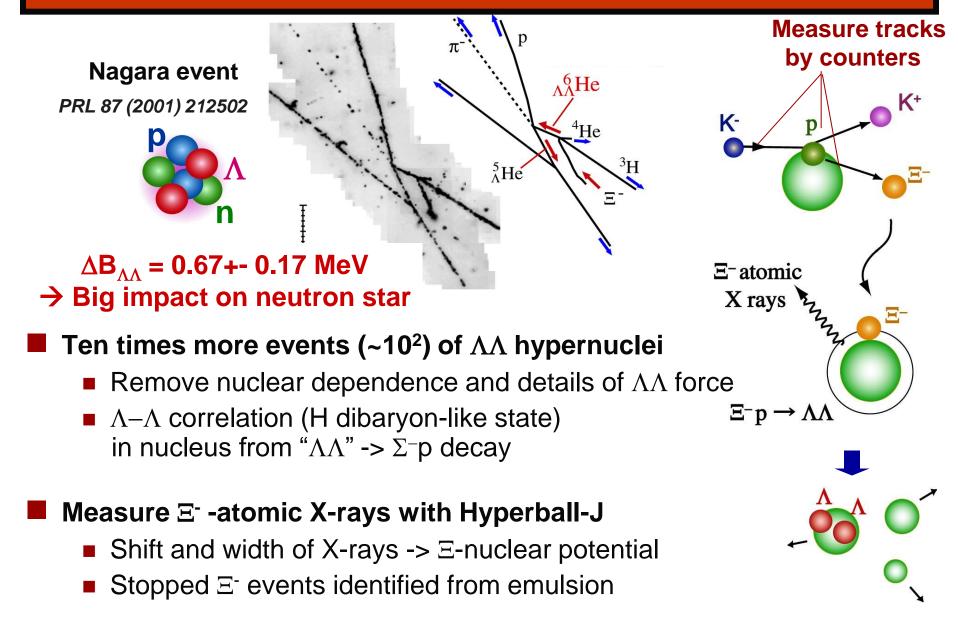
<u>ΛΛ hypernuclei via</u>

Nakazawa (Hyp-X conf.)

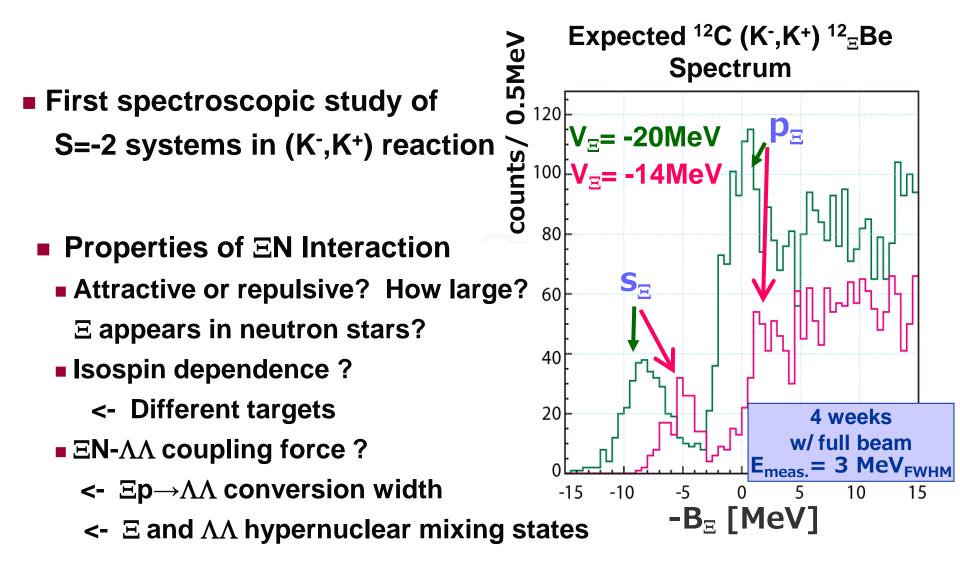
emulsion+counter hybrid method (KEK E373)



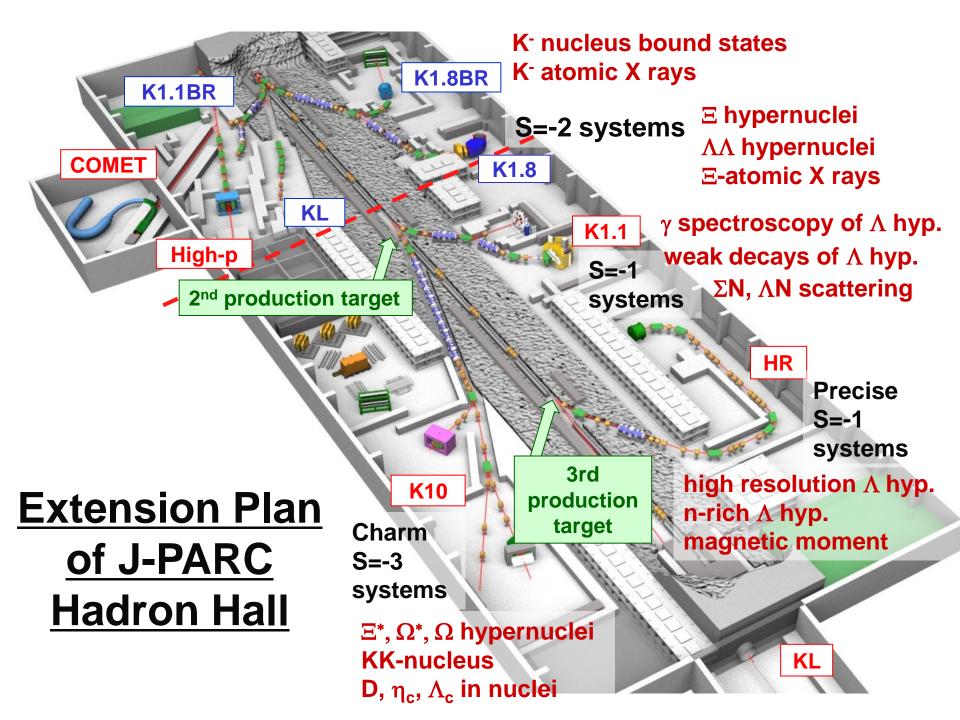
J-PARC E07 (Nakazawa, Imai, Tamura et al.) S=-2 Systems with Emulsion-Counter Hybrid Method



J-PARC E05 (Nagae et al.) $K^-p \rightarrow \Xi^-K^+$ Ξ -hypernuclear spectroscopy by (K^-, K^+)



5. Future Plans at J-PARC



6. Summary

- Strangeness nuclear physics studies BB forces, impurity effects, inmedium baryon properties and provides clues for neutron star matter.
- A hypernuclear spectroscopy: Neutron-rich ${}^{10}_{\Lambda}$ Li and ${}^{6}_{\Lambda}$ H observed. Will investigate Λ NN force in n-rich environment.
- γ-spectroscopy of Λ hypernuclei:
 From p-shell data, spin-dependent ΛN int. strengths determined.
 In-medium g_Λ to be measured from B(M1).
- Σ-nuclear systems: strongly repulsive potential observed.
 Σ-p scattering experiment planned to confirm the quark Pauli effect.
- Double strange systems: ${}^{6}_{\Lambda\Lambda}$ He revealed $\Lambda\Lambda$ force weakly attractive. Ξ hypernuclei to be first studied to measure the Ξ N force.
- Extension of J-PARC Hadron Hall will extend strangeness nuclear physics further.

Collaborations with theoretical studies of lattice BB forces, nuclear structure, etc. are more and more important.

2012-2016 新学術領域 「実験と観測で解き明かす中性子星の核物質」

Grant-in-aid for innovative area: "Nuclear Matter in neutron Stars investigated by experiments and astronomical observations"

> Project leader: H. Tamura Theory group: A. Ohnishi