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Nucleosynthesis in Supernovae and GRBs and Neutrino Oscillation

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COSNAP-Collaboration

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PURPOSE

Study SUPERNOVA-*v* and **CMB-Anisotropies:**

1. to constrain the v–oscillation parameters θ_{13} and Δm_{13}^2 . 2. to constrain the v–total mass.

Various roles of v's in SN-nucleosynthesis



Neutrino-driven relativistic jet (Harikae + '09, '10



- Special relativistic MHD code (Takiwaki et al 2009)
- Neutrino heating is calculated by ray-tracing in flat timespace.
- ◆ Realistic EOS of Shen et al (1998) is implemented.
- Initial data is taken from 35OC model (Woosley&Heger 2006)





R-process in Pair v-Heated Collapsar Model for GRB

K. Nakamura, S. Sato, S. Harikae, T. Kajino and G.J. Mathews (2011), submitted to ApJ.

Neutron-rich condition for successful r-process:



relative abundance

Tantalum(^{180,181}Ta) ¹³⁸La

¹⁸¹Ta_g(stable), ¹⁸⁰Ta_g(unstable, $\tau_{1/2} = 8h$), ¹⁸⁰Ta^m(isomer, $\tau_{1/2} > 10^{15}y$)

¹⁸⁰Ta is the rarest isotope in the Solar–Systerm and even in the Universe!

Where was ¹⁸⁰Ta synthesized ?



¹⁸⁰Ta-genesis needs Quantum Phys. + SN Hydro-dyn.

Solar-¹⁸⁰Ta is all "**ISOMER**" with $T_{1/2} > 10^{15}$ y!

- Long lived ¹⁸⁰Ta^m is excited in hot SN-photon bath.
- Intermediate states are depopulated to the ground state, which decays in 8 hours.



Result from v-Nucleosynthesis

T. Hayakawa, T. Kajino, S. Chiba, and G.J. Mathews, Phys. Rev. C81 (2010), 052801®



About 40% ¹⁸⁰Ta^m survives in supernova explosion.

Then, both ¹³⁸La and ¹⁸⁰Ta abundances can be consistently reproduced by the CC-int. of v_e and $\overline{v_e}$ of

Oscillation (MSW) Effect on Supernova v-Process

SN II: Yoshida, Kajino & Hartman, Phys. Rev. Lett. 94 (2005), 231101. SNIc + II: Nakamura, Yoshida, Shigeyama, Kajino, ApJL 718 (2010), L137.



Our Theoretical Prediction

⁷Li/¹¹B-Ratio

Yoshida, Kajino et al . 2005, PRL94, 231101; 2006, PRL 96, 091101; 2006, ApJ 649, 319; 2008, ApJ 686, 448.

Astrophysics:

Mass Hierarchy

 Δm_{13}^2

13-Mixing Angle

 θ_{13}

Long Baseline Exp:

• MINOS

Daya Bay

T2K (Kamioka)

June 14, 2011

July 29, 2011

RENO (KOREA)

Double CHOOZ



What is the lower limit to $sin^2 2\theta_{13}$?



	Selected for a Viewpoint in <i>Physics</i>	weak anding	
PRL 107, 041801 (2011)	PHYSICAL REVIEW LETTERS	22 JULY 2011	

Indication of Electron Neutrino Appearance from an Accelerator-Produced Off-Axis Muon Neutrino Beam

Compassison between MINOS and T2K



L. Whitehead, BNL

Our Theoretical Prediction

⁷Li/¹¹B-Ratio

Yoshida, Kajino et al . 2005, PRL94, 231101; 2006, PRL 96, 091101; 2006, ApJ 649, 319; 2008, ApJ 686, 448.





THE ASTROPHYSICAL JOURNAL LETTERS, 730:L7 (5pp), 2011 March 20

HINTS FOR NEUTRINO-PROCESS BORON IN PRESOLAR SILICON CARBIDE GRAINS FROM SUPERNOVAE

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Mass Hierarchy, Normal or Inverted ?

Mathews, Kajino, Aoki and Fujiya, arXiv:1108.0725 (2011).



v-Asymmetry under the Strong Dipole (Poroidal) Magnetic Field

Fundamental Interactions among Hadrons (p, n, Λ , Σ ...) and Lepton (e,v...) at High- ρ and High-T in Relativistic Field Theory and QCD

Maruyama, Kajino, Yasutake, Cheoun, & Ryu, PRD83 (2011), 081302 (R).





Neutrino scattering and absorption process inside the magnetized Neutron star (10¹⁵G) is asymmetric.

 \Rightarrow ~ 2 % asymmetric v-abs. (drift)

⇒ Enough for Pulsar-Kick ~500km/s !

Magnetic Field and v-interactions

★ Poroidal → Pulsar Kick
 ★ Toroidal → Twist Mode → Spin down & growth of intsability

T. Kuroda, with Yasutake, Maruyama, Hidaka & Kajino

Initially dipole → **Double Toroidals**

Initially no dipole \rightarrow Single Toroidal





Why Amino Acids on the Earth, All Left-Handed?

Chitrality, earth origin or universal?



- ★ Neutrinos are all left-handed!
- ★ Supernovae with strongly magnetized neutron star or BH emit intensive flux of neutrinos over 10¹⁰ yrs!
- ★ SN ejecta including ¹⁴N interact with neutrino under strong magnetic field!
- ★ Neutrino-¹⁴N coupling is asymmetric & chiral selective!

Boyd, Kajino, & Onaka suggested that the L-handed chirality of amino acids is UNIVERSAL! (Astrobio. 10 (2010), 561-568; Int. J. Mol. Sci. 12 (2011), 3432)

Magnetized Supernovae



¹⁴N lives !



Mann and Primakoff (Origins of Life, 11 (1981), 255) suggested β -decay of 14C, but it's too SLOW!

CONCLUSION

We propose a new astrophysical method to determine the unknown v-oscillation parameters, θ_{13} and mass hierarchy Δm_{13}^2 , simultaneously in terms of the supernova v-process nucleosynthesis of ¹⁸⁰Ta, ¹³⁸La, ⁷Li and ¹¹B By taking account of the MSW effects.

Combining the recent detection of ⁷Li/¹¹B isotopic ratio in presolar X-grains and the T2K + MINOS results of long-baseline neutrino oscillation experiments on θ_{13} , we can conclude that the "inverted mass hierarchy" is more preferred.

Collaboration, highly desirable in Theory plus Exp/Obs. as required/demoinstrated in SN-neutrino project! Nucleosynthesis—c.c. SNe—v—Magnetic Field—Life