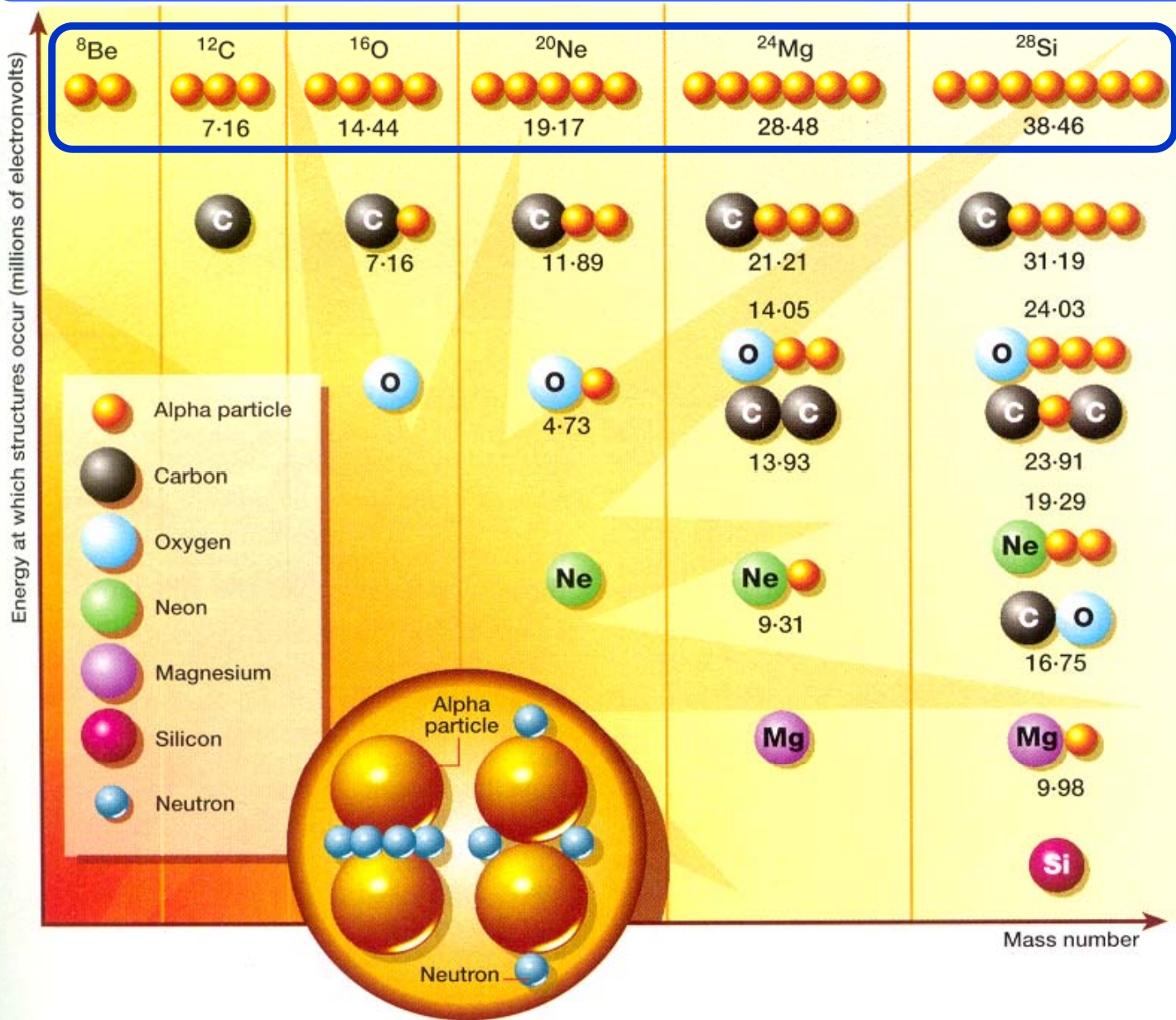


原子核における α 粒子凝縮状態の研究

Yasuro Funaki (Hiyama lab., RIKEN)

素核宇融合による計算基礎物理学の進展
—マイクロとマクロのかけ橋の構築—
@合歓の里, 2011年12月3日—5日.

Prediction of cluster states in light nuclei (Ikeda Diagram)



The most tightly bound light cluster

α particle (quartet)

$E/A \sim 1 \text{ MeV}$ $E^* \sim 20 \text{ MeV}$
stiff

The most elemental subunit in nuclear cluster structures.

Pair (deuteron) is less bound in nuclear system.

$E/A \sim 1 \text{ MeV}$

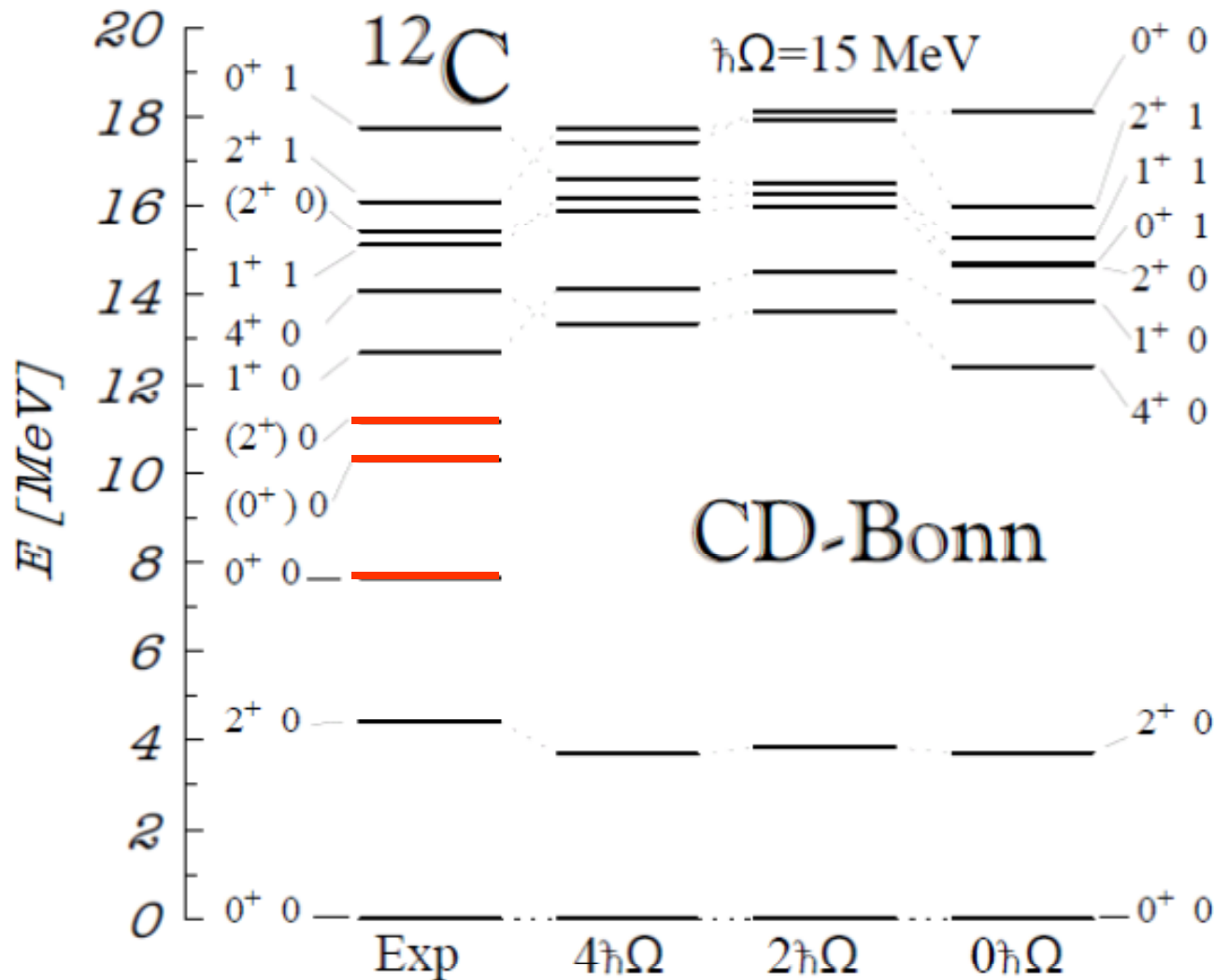
Classified according to the Threshold Rule.

K. Ikeda et al, PTP suppl. Extra num., 464 (1968).

Typical mysterious 0^+ states in nuclear structure problem

0_2^+ state of ^{12}C (Hoyle state) **indispensable to ^{12}C production in stars**

Ab initio non-core shell model calculation



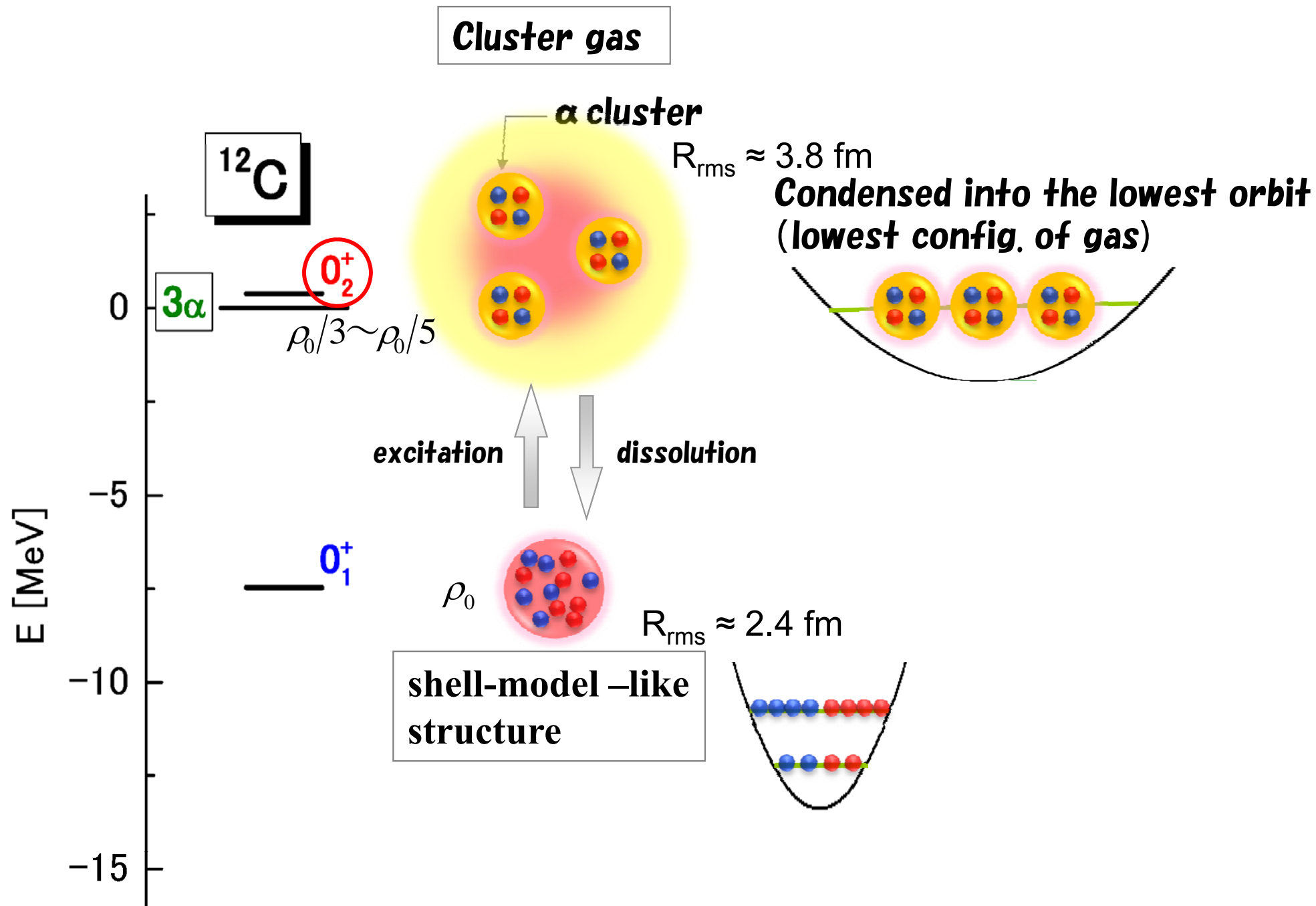
0_2^+ , 2_2^+ , 0_3^+ states : missing

0_2^+ state: excitation energy is not lower than 20 MeV

The typical excited states which resist a shell model description

P. Navratil et al., PRL 84, 5728 (2001).

First example of alpha cond. state



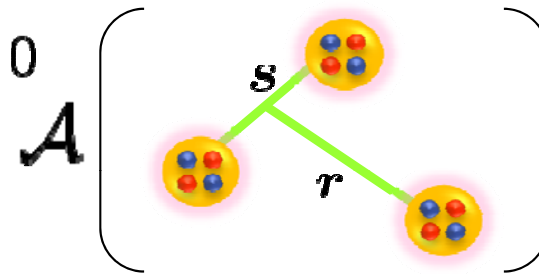
First example of α condensate state in finite nuclei

RGM (Full 3α) vs 3α cond. (3α confined in $0S$ orbit)

\mathcal{A} : antisymmetrizer acting on 12 nucleons

$$\langle \phi^3(\alpha) | H - E | \mathcal{A}[\chi(\mathbf{s}, \mathbf{r}) \phi^3(\alpha)] \rangle = 0$$

M. Kamimura, NPA 351, 456 (1981).



	Exp.	RGM
Energy (MeV)	7.65	7.74
α decay width (eV)	8.7 ± 2.7	7.7
$M(0_2^+ \rightarrow 0_1^+)$ (fm^2)	5.4 ± 0.2	6.7
$B(E2: 0_2^+ \rightarrow 2_1^+)$ ($\text{e}^2 \text{fm}^4$)	13 ± 4	5.6

First example of α condensate state in finite nuclei

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The Solution of 3α RGM eq. of motion is almost equivalent to the 3α cond. w.f.
 The full 3α problem gives the 3α condensate w.f. as its solution!

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M. Kamimura, NPA 351, 456 (1981).



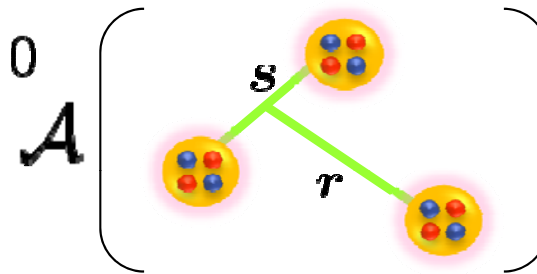
$$\chi = \prod_{i=1}^3 \exp\left(-\frac{2}{B^2}(\mathbf{X}_i - \mathbf{X}_G)^2\right)$$

X_i : com coordinate of the i -th α \mathcal{A}

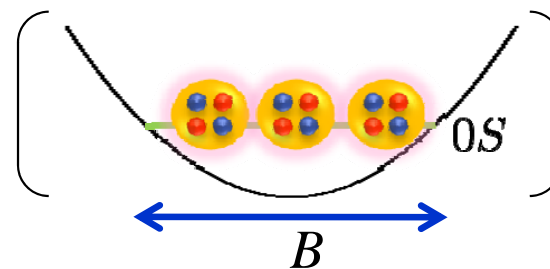
X_G : total com coordinate

Y. F. et al., PRC 67, 051306(R) (2003).

RGM w.f.



3α cond. w.f.



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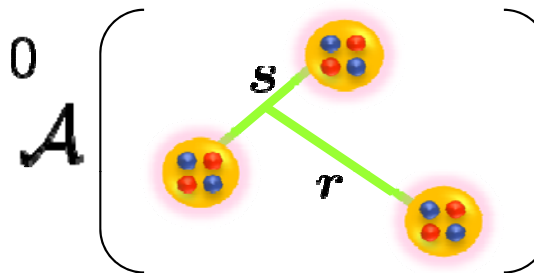
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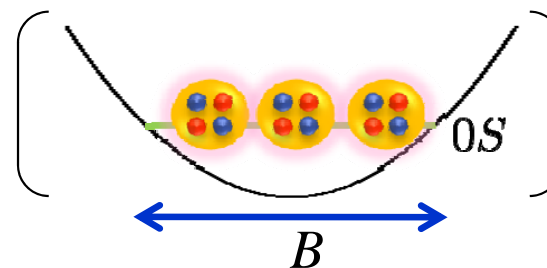
X_G : total com coordinate

Y. F. et al., PRC 67, 051306(R) (2003).

RGM w.f.



3α cond. w.f.



3α clustering also appears starting without assumption of α 's by FMD & AMD

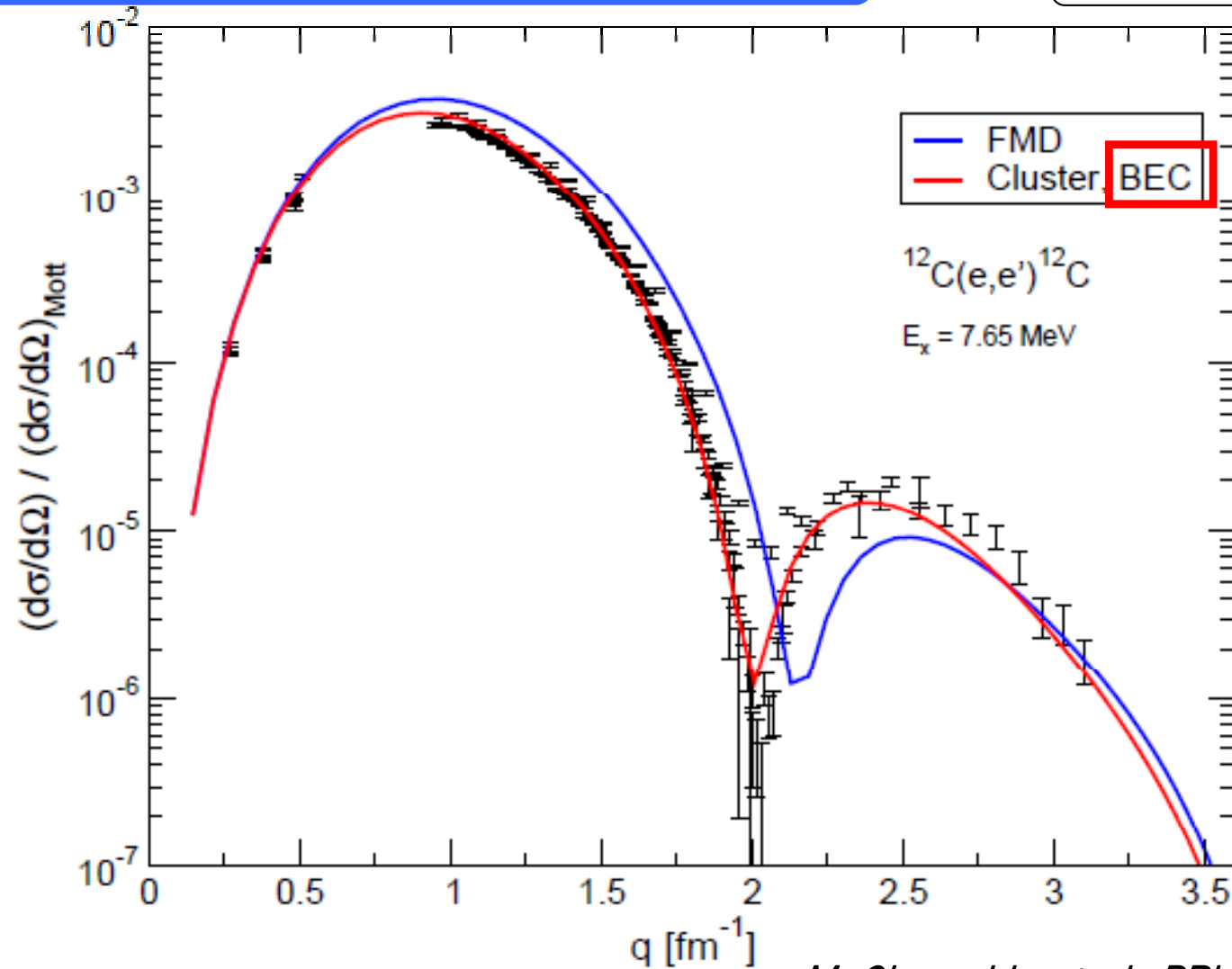
M. Chernykh, T. Neff et al., PRL 94, 032501 (2007).

Y. Kanada-En'yo, PTP 117, 655 (2007).



Electron Scattering Data ($0_1^+ \rightarrow 0_2^+$)

© T. Neff



M. Chernykh. et al., PRL 98, 032501 (2007)

Very nice reproduction by THSR w.f. (BEC)

“BEC” from Y.F. et al., EPJA 28, 259(2006)

Direct information of alpha condensation for the Hoyle state

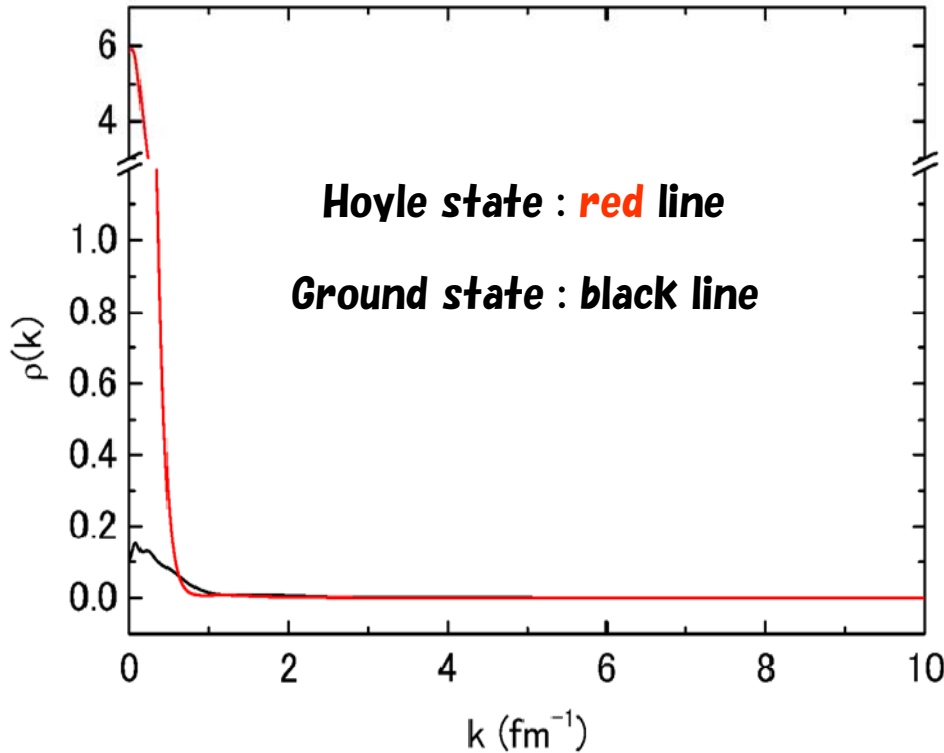
via 3α OCM (Orthogonality Condition Model)

$$\rho(k) = \int dr dr' \frac{e^{-ik \cdot r}}{(2\pi)^{3/2}} \rho(r, r') \frac{e^{ik \cdot r'}}{(2\pi)^{3/2}}$$

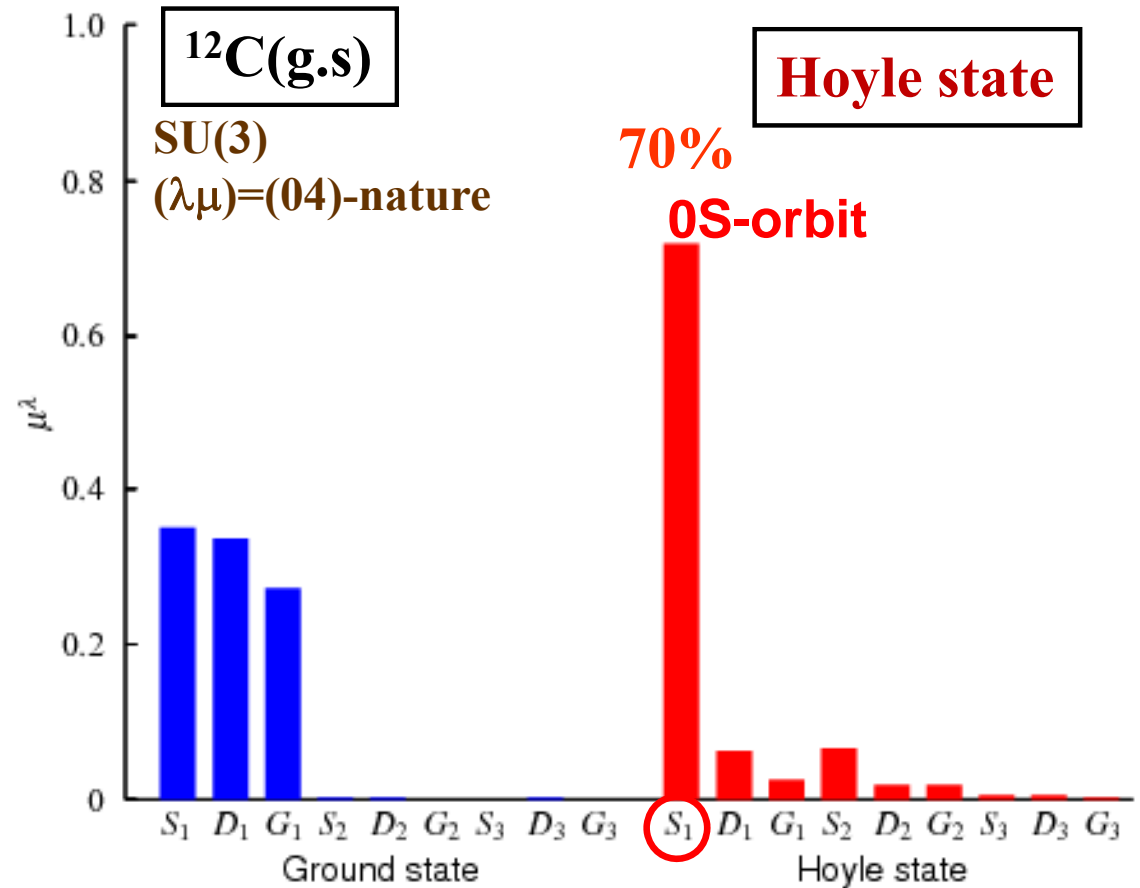
$$\int dr' \rho(r, r') \phi(r') = \mu \phi(r)$$

Momentum distribution of α -particle

Occupation probability of single α -orbit



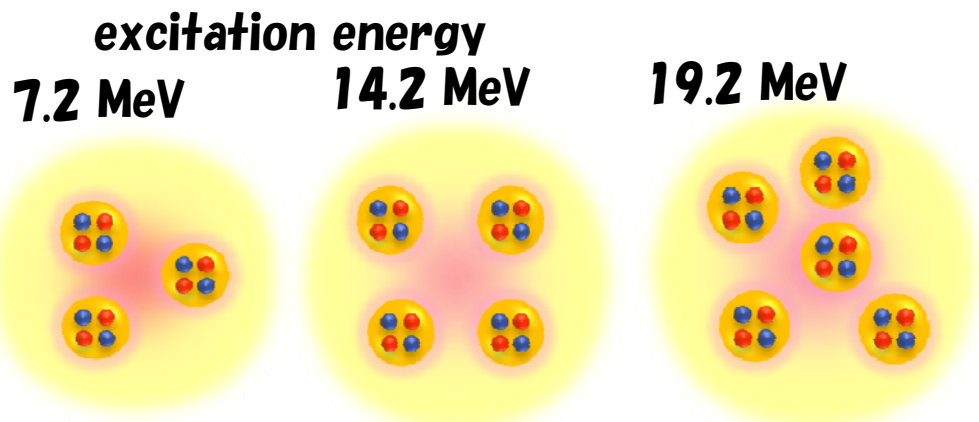
δ function-like peak around zero momentum



T. Yamada and P. Schuck, EPJA 26, 185 (2005).

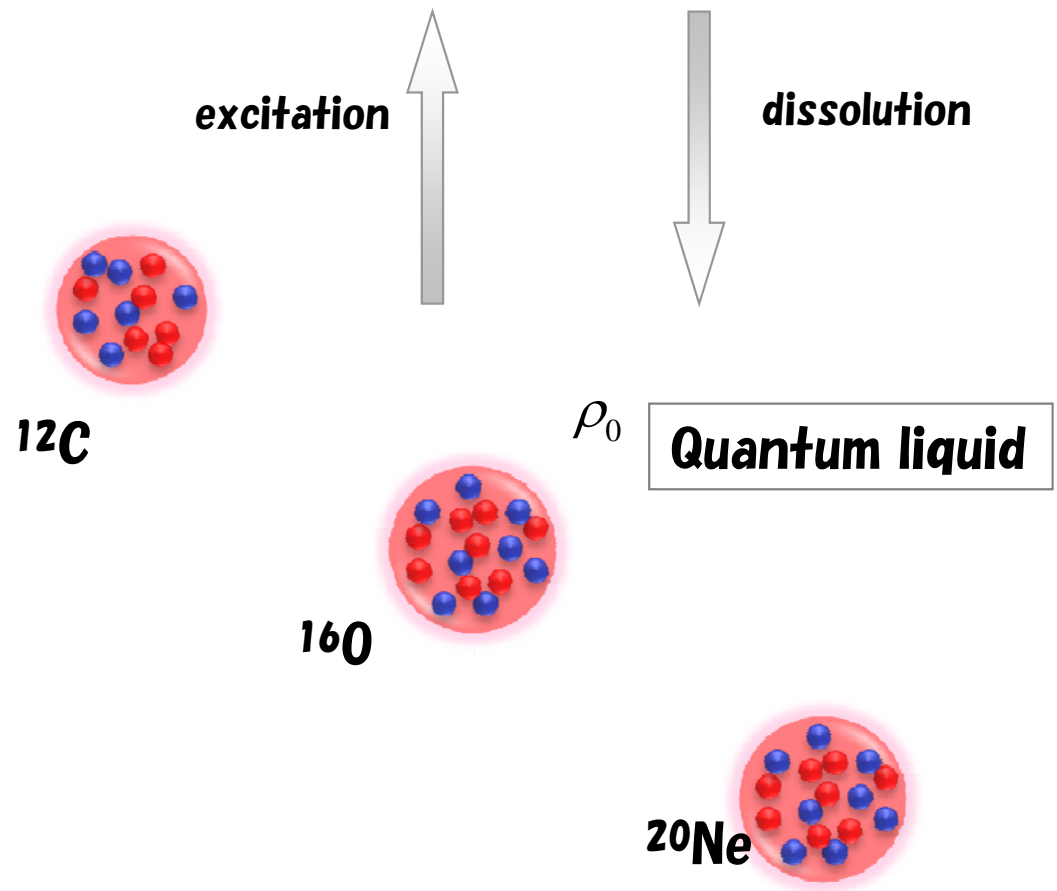
“gas phase” in finite nuclei

Energy



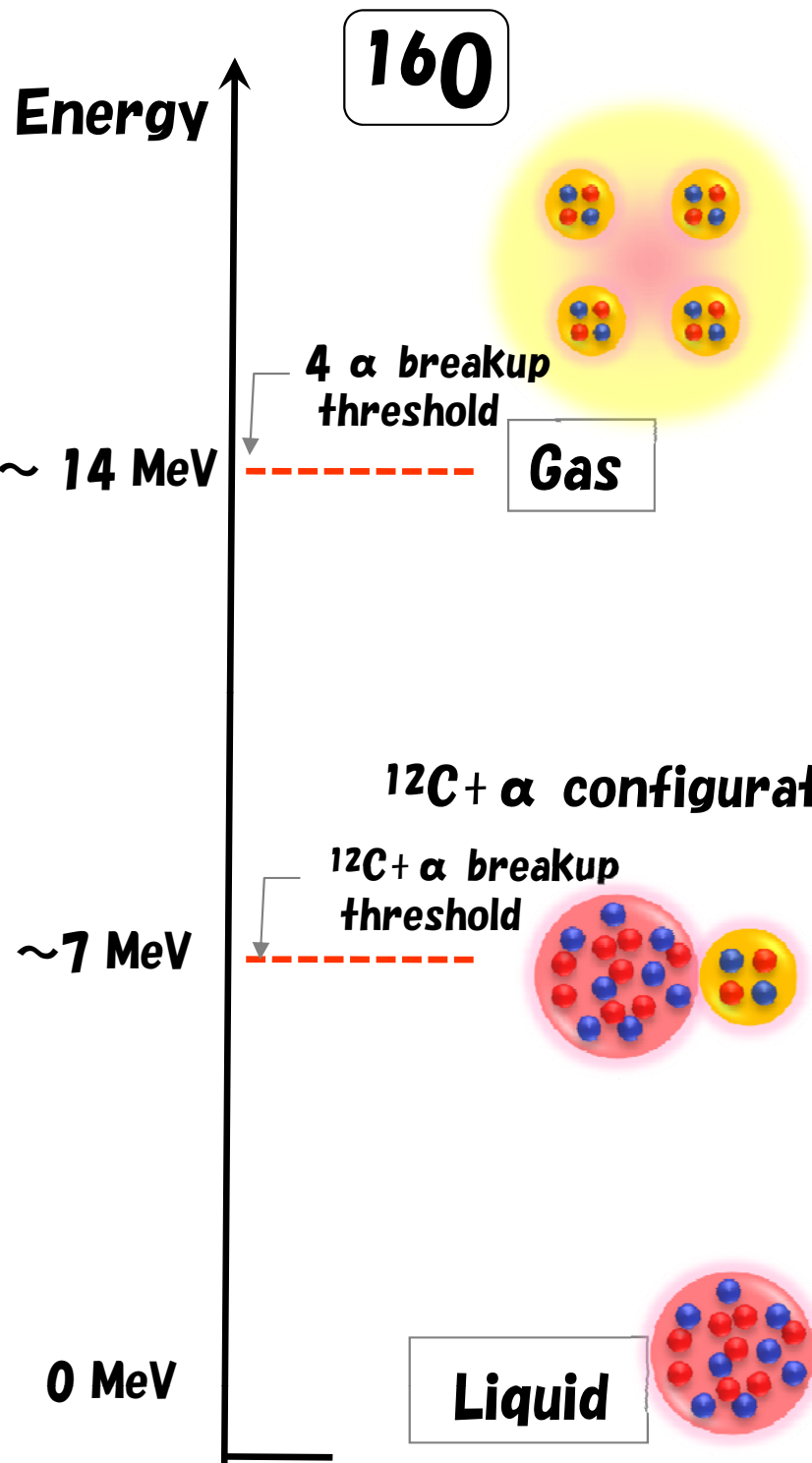
**Infinite nuclear matter
(low density)
 $< \rho_0/5$
Crust of neutron star?**

$n\alpha$ threshold energy $\rho_0/3 \sim \rho_0/5$ **Cluster gas**



**Investigation in
heavier nuclei
than ^{12}C**

Analogue to the Hoyle state in ^{16}O ?



To describe 4 α cluster states

× mean field model

× $^{12}\text{C} + \alpha$ cluster model

○ 4 α cluster model

Very important : simultaneous description of lower states (g.s., $^{12}\text{C} + \alpha$)

Fully solving 4 α - particles relative motions (4 α OCM)

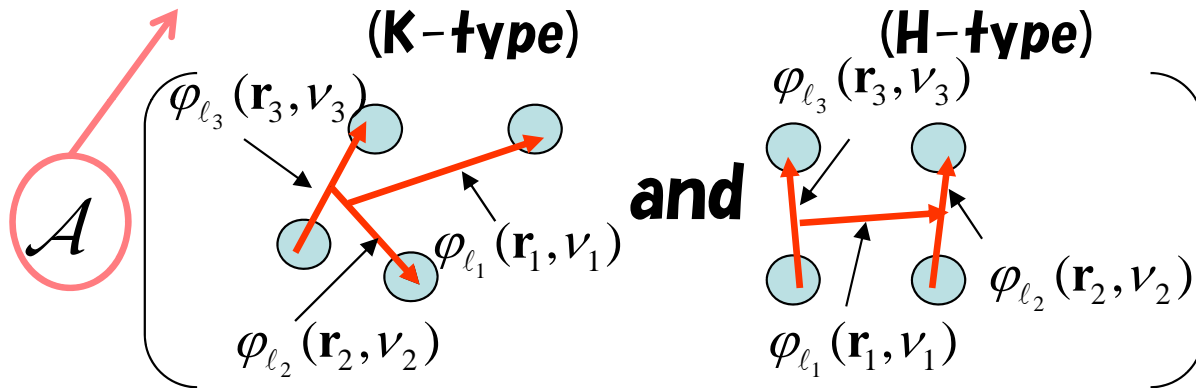
Present: Larger model space

$$\varphi_{\ell m}(\mathbf{r}, \nu) = N_{\ell}(\nu) r^{\ell} \exp(-\nu r^2) Y_{\ell m}(\mathbf{r})$$

Gaussian basis (GEM)

E. Hiyama et al. Prog. Part. Phys. 51, 223(2003).

Approximately taken into account



**Adopted angular momentum channels: $[[l_1, l_2], l_3]$ ($l_3 + l_2 + l_1 \leq 8$) (up to now, ≤ 5)
Including $l_3, l_2, l_1 = 4$**

Total w.f.

$$\Psi_{\text{OCM}}(J_k^{\pi}) = \sum_{\{l\}\{\nu\}} A_{l_1, l_2, l_{12}, l_3}^{(k)}(\nu_1, \nu_2, \nu_3) \hat{S} \left[\left[\varphi_{l_1}(\mathbf{r}_1, \nu_1), \varphi_{l_2}(\mathbf{r}_2, \nu_2) \right]_{l_{12}}, \varphi_{l_3}(\mathbf{r}_3, \nu_3) \right]_J$$

$A_{l_1, l_2, l_{12}, l_3}^{(k)}(\nu_1, \nu_2, \nu_3)$: Determined by diagonalizing Hamiltonian

Hamiltonian of 4 α OCM

$$H = T + \sum_{i < j} \left[V_{2\alpha}(r_{ij}) + V_{2\alpha}^{Coul}(r_{ij}) \right] + V_{3\alpha} + V_{4\alpha} + V_{Pauli}$$

Pauli blocking operator on $\alpha - \alpha$ motions

$$V_{Pauli} = \lim_{\lambda \rightarrow \infty} \lambda \sum_{2n+\ell < 4} \sum_{ij} |u_{n\ell}(r_{ij})\rangle \langle u_{n\ell}(r_{ij})|$$

Pauli forbidden state: h.o.w.f.

2-body force (folding MHN force)

$$V_{2\alpha}(r) = \sum_n V_n^{(2)} \exp(-\beta_n^{(2)} r^2)$$

Coulomb force

$$V_{2\alpha}^{Coul}(r) = \frac{4e^2}{r} \text{erf}(ar)$$

Phenomenological 3-body force (repulsive)

$$V_{3\alpha} = V^{(3)} \sum_{i < j < k} \exp[-\beta(r_{ij}^2 + r_{jk}^2 + r_{ki}^2)]$$

$$V^{(3)} = 87.5 \text{ MeV}, \quad \beta = 0.15 \text{ fm}^{-2}$$

Phenomenological 4-body force (repulsive)

$$V_{4\alpha} = V^{(4)} \exp[-\beta(r_{12}^2 + r_{13}^2 + r_{14}^2 + r_{23}^2 + r_{24}^2 + r_{34}^2)]$$

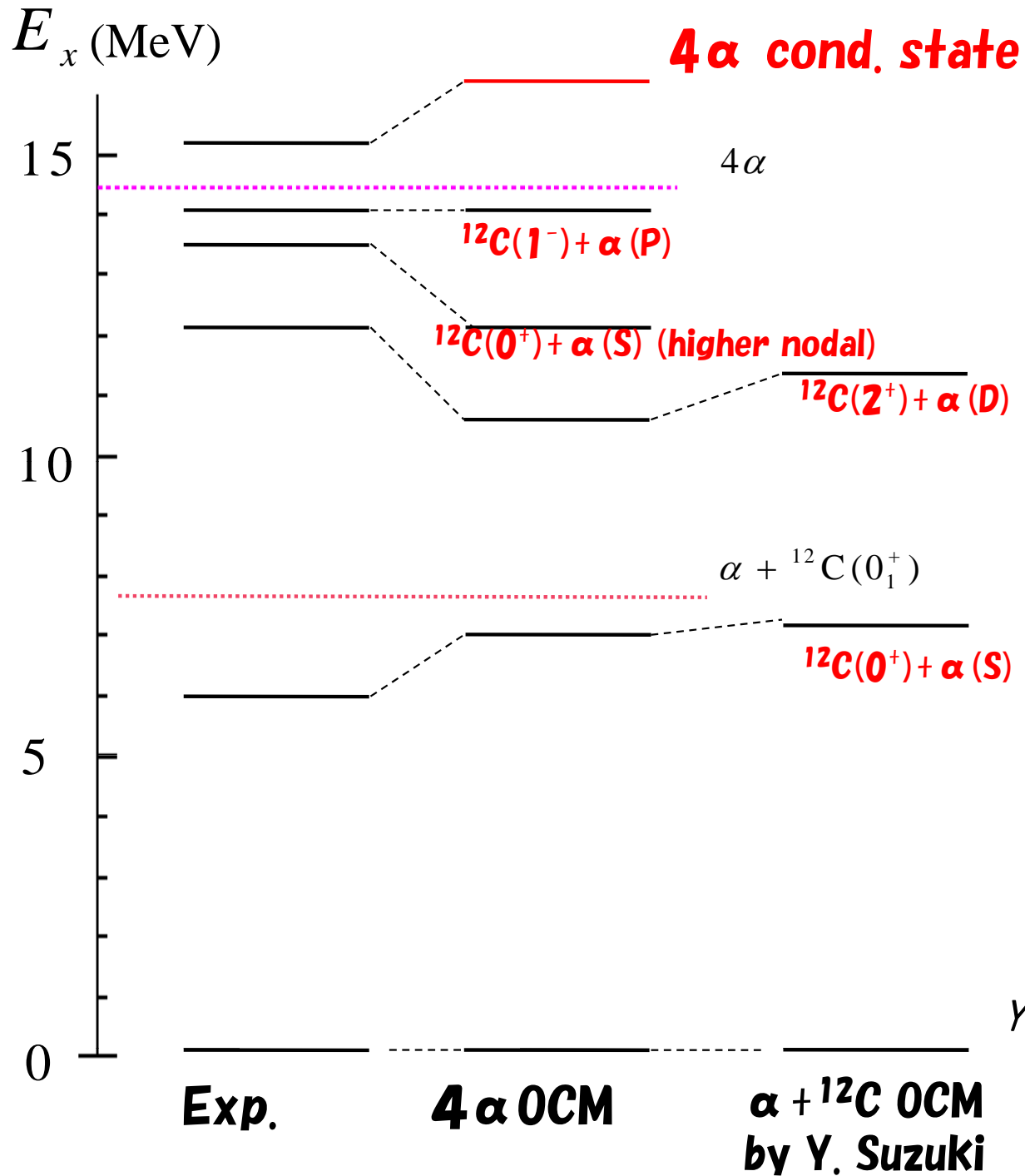
$$V^{(4)} = 12000 \text{ MeV}, \quad \beta = 0.15 \text{ fm}^{-2}$$

Energies from 4 α threshold

	Cal. (MeV)	Exp. (MeV)
$^{12}\text{C}(\text{g.s.})$	<u>-7.32</u>	<u>-7.28</u>
$^{12}\text{C}(2_1^+)$	-4.88	-2.84
$^{12}\text{C}(4_1^+)$	2.06	6.43
$^{12}\text{C}(0_2^+)$	0.70	0.38
$^{16}\text{O}(\text{g.s.})$	<u>-14.2</u>	<u>-14.44</u>

$$|\langle V_{3\alpha} \rangle|, |\langle V_{4\alpha} \rangle| < \frac{7}{100} |\langle V_{2\alpha} \rangle|$$

0^+ spectra, rms radii, monopole matrix elements



0_4^+ state: T. Wakasa, Y. F. et al.,
PLB 653, 173 (2007).

Y. F. et al., PRL 101, 081502 (2008).

0^+ spectra, rms radii, monopole matrix elements**Large monopole matrix element can be the evidence of cluster states.***T. Yamada, Y. F. et al., PTP120, 1139 (2008).*

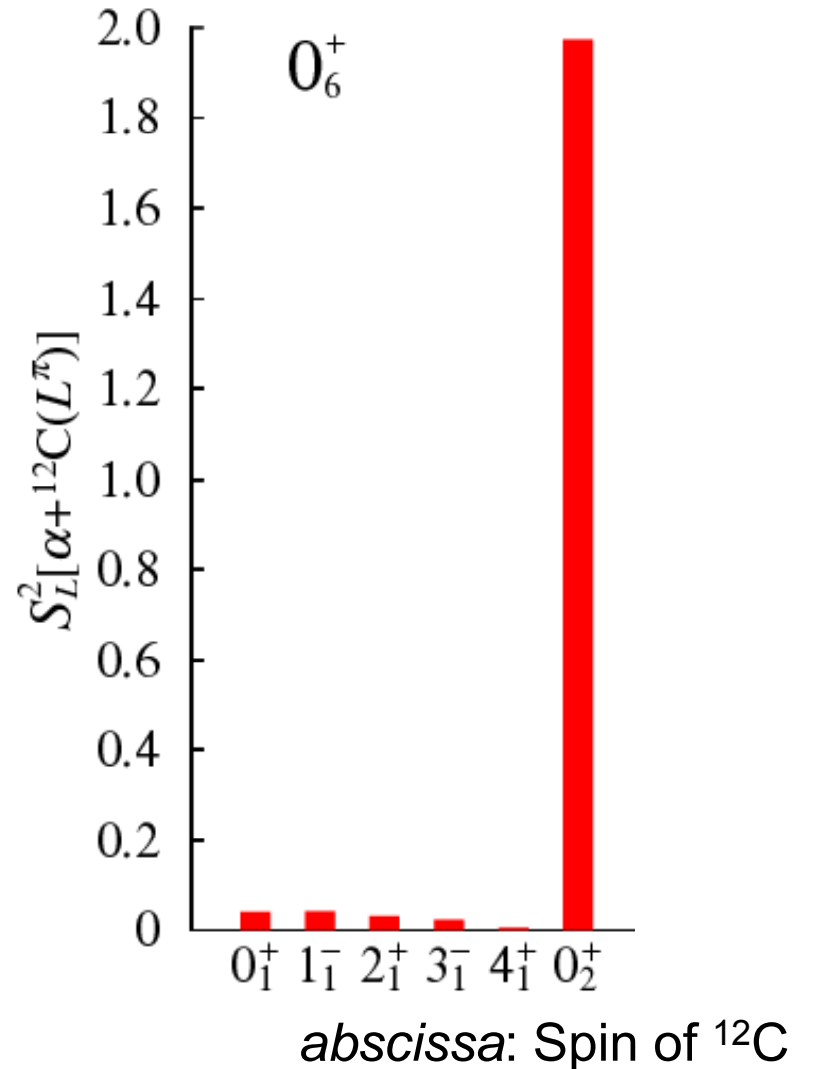
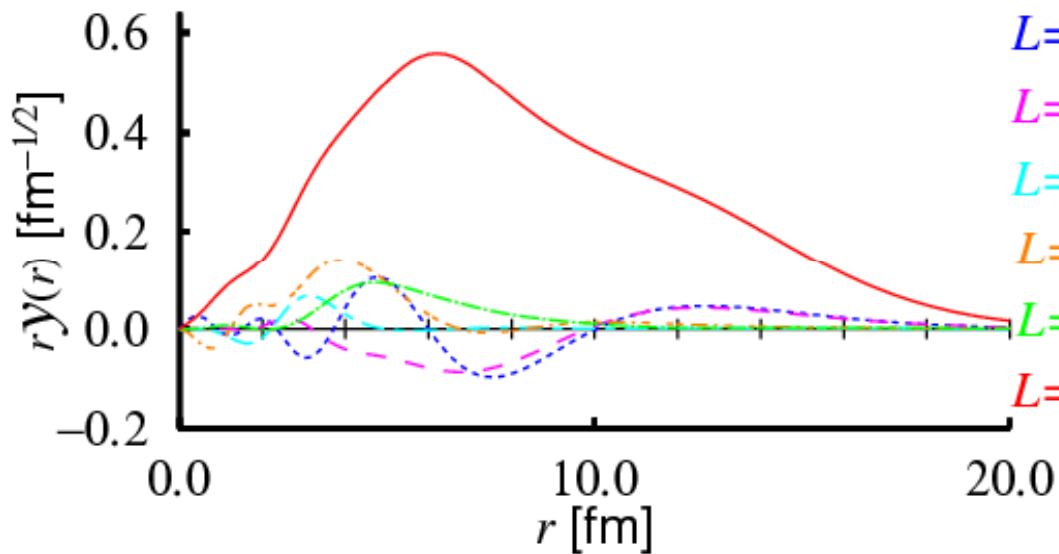
	E_x [MeV]	Experimental data			4α OCM		
		R [fm]	M(E0) [fm ²]	Γ [MeV]	R [fm]	M(E0) [fm ²]	Γ [MeV]
0^+_1	0.00	2.71			2.7		
0^+_2	6.05		3.55		3.0	3.9	
0^+_3	12.1		4.03		3.1	2.4	
0^+_4	13.6		no data	0.6	4.0	2.4	0.60
0^+_5	14.0		3.3	0.185	3.1	2.6	0.20
0^+_6	15.1		no data	0.166	5.6	1.0	0.14

over 15%
of total EWSR20%
of total EWSR

S-factor : $^{12}\text{C} + \alpha$ and $^8\text{Be} + ^8\text{Be}$ components for the 0_6^+ state

$$r \times \mathcal{Y}_{IL, J=0}(r) = r \times \sqrt{\frac{4!}{3!1!}} \left\langle \left[\frac{\delta(r-r')}{rr'} Y_L(\hat{r}') \Psi_{\text{OCM}}(^{12}\text{C}(I)) \right]_0 \middle| \Psi_{\text{OCM}}(0_k^+) \right\rangle$$

$$S_{IL}^2(J=0) = \int dr \left(r \times \mathcal{Y}_{IL, J=0}(r) \right)^2$$



Analogue to the Hoyle state

$\alpha + ^{12}\text{C}$ (Hoyle) configuration is dominant.
 ^{12}C (Hoyle): 3 α condensate



4 α condensate

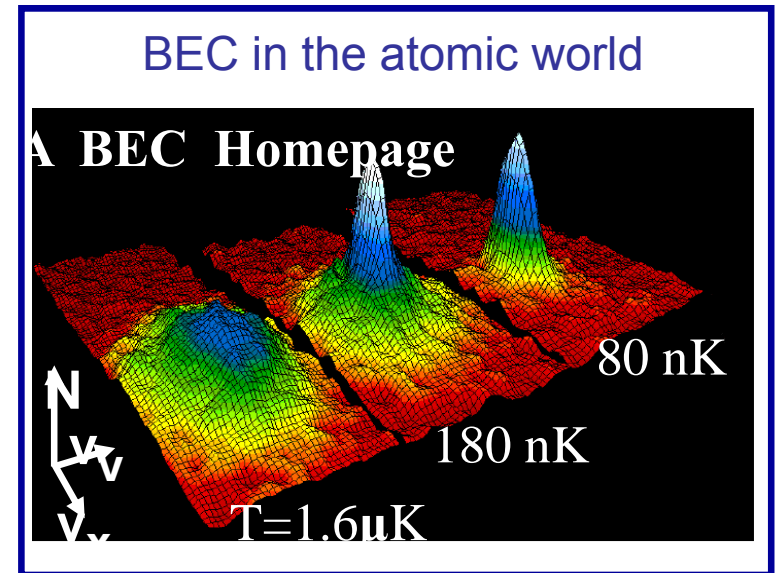
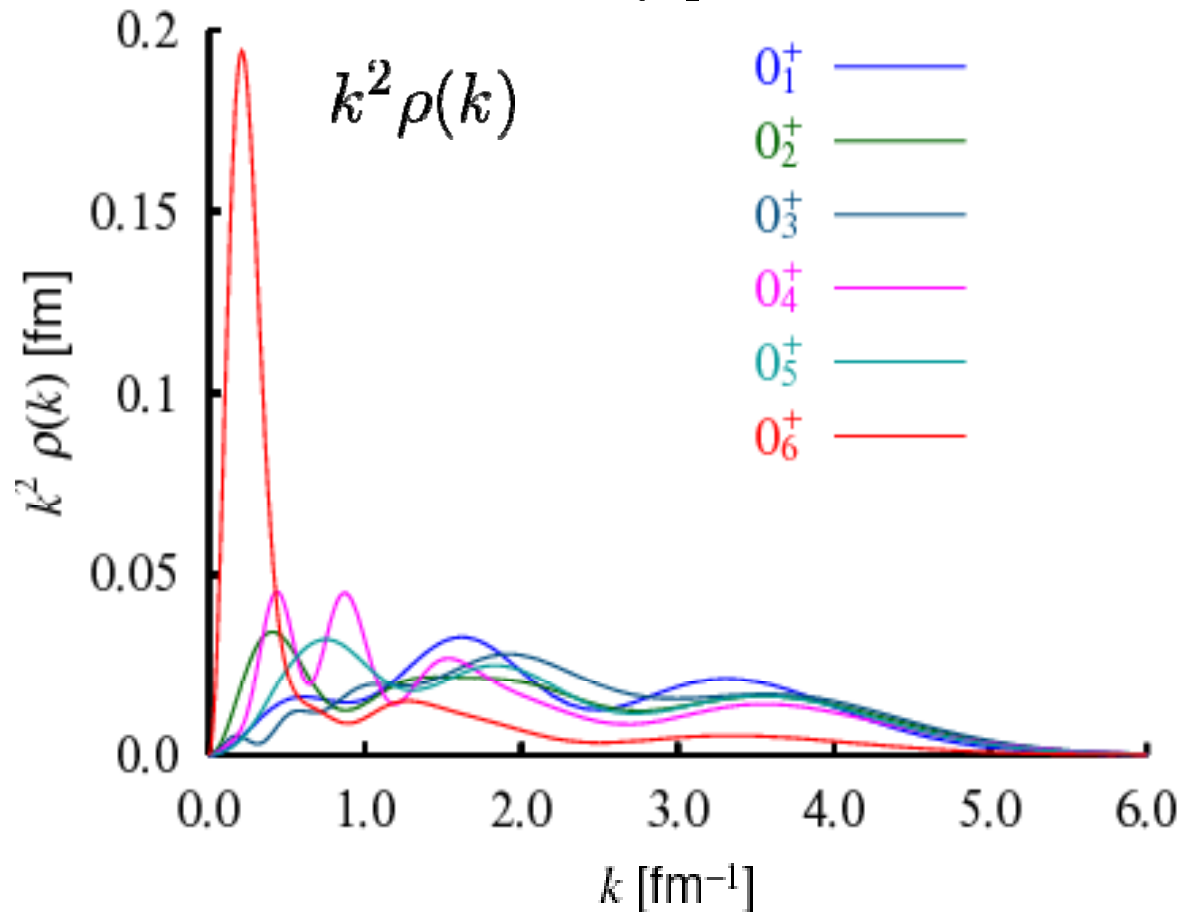
Momentum distributions of the α particles

$$\rho(k) = \int d\mathbf{r} d\mathbf{r}' \frac{e^{-i\mathbf{k}\cdot\mathbf{r}}}{(2\pi)^{3/2}} \rho(\mathbf{r}, \mathbf{r}') \frac{e^{i\mathbf{k}\cdot\mathbf{r}'}}{(2\pi)^{3/2}}$$

$$\rho(\mathbf{r}, \mathbf{r}') = \frac{1}{4} \sum_{i=1}^4 \langle \Psi_{\text{OCM}}(0_k^+) | \delta(\mathbf{r}_i - \mathbf{X}_G - \mathbf{r}') \rangle \langle \delta(\mathbf{r}_i - \mathbf{X}_G - \mathbf{r}) | \Psi_{\text{OCM}}(0_k^+) \rangle$$

r_i : coordinate of the i -th α particle

X_G : coordinate of total center-of-mass

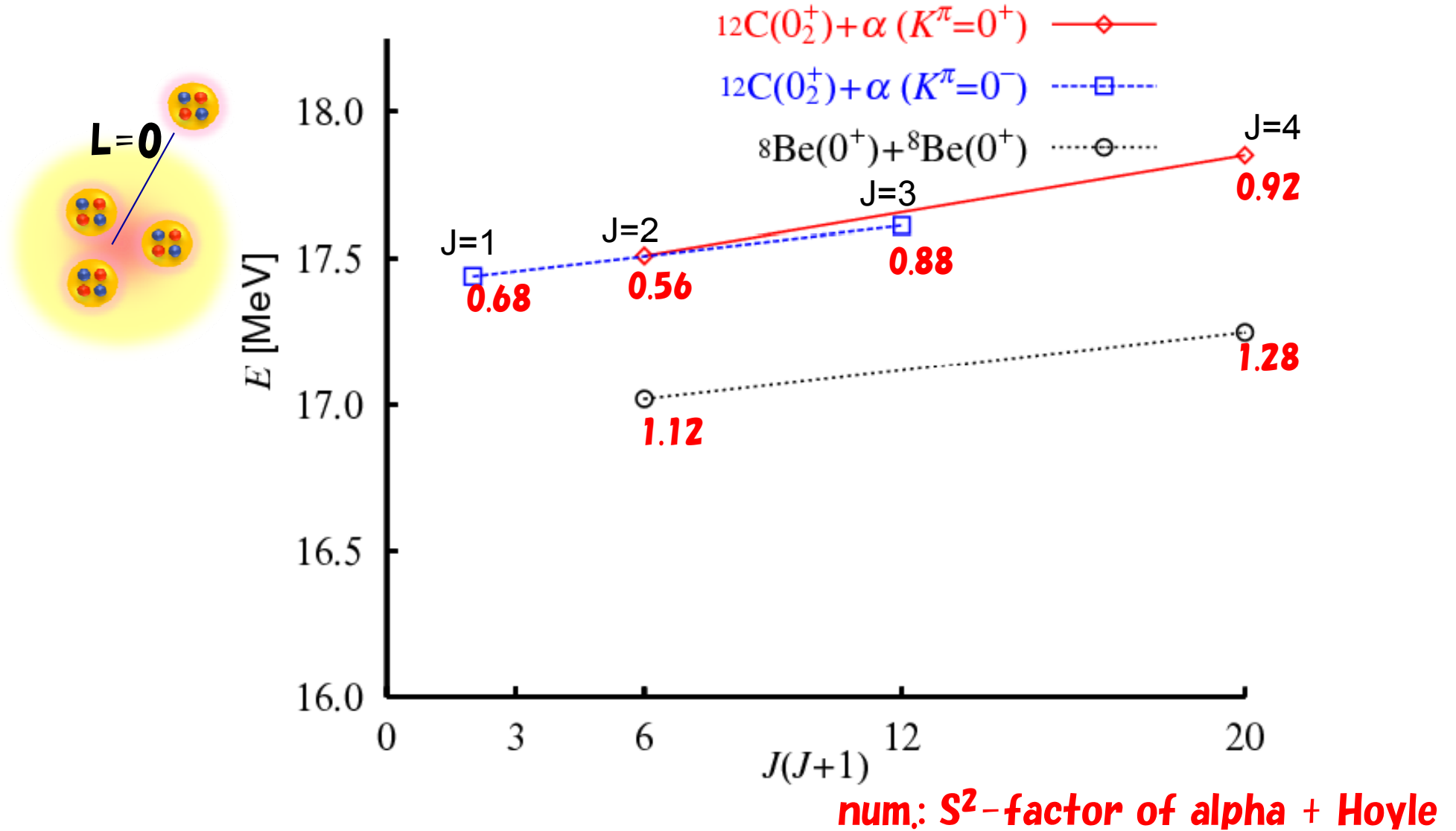


0_6^+ : delta-function-like peak at zero momentum

4 α condensate state character. de Broglie w.l. $\lambda = \frac{2\pi}{\sqrt{\langle k^2 \rangle}} \geq 20$ fm

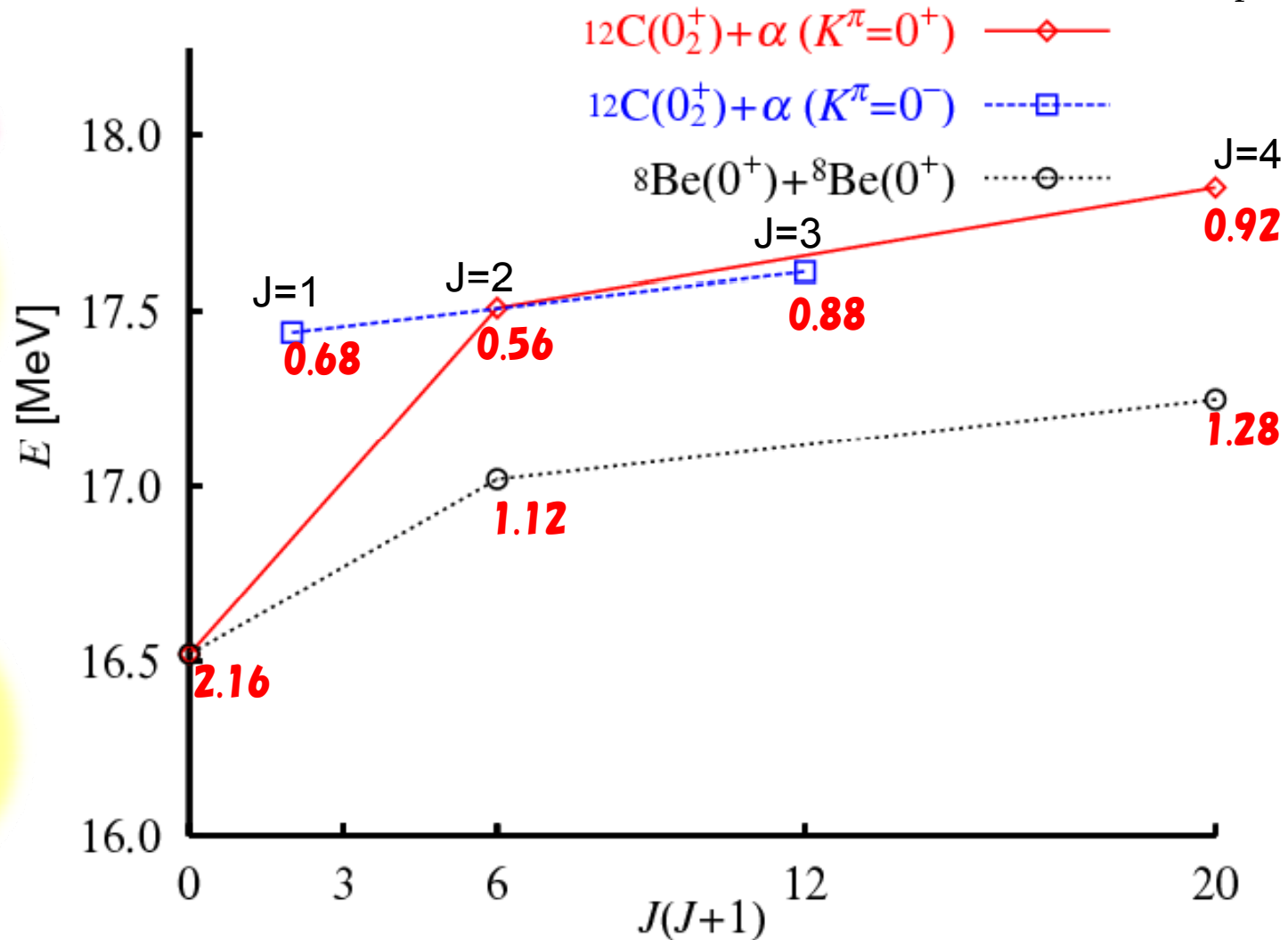
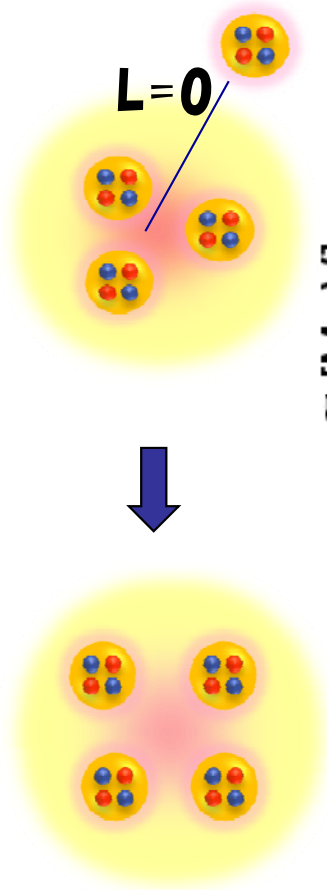
Rotational bands of Hoyle + alpha, ${}^8\text{Be}+{}^8\text{Be}$?

$$E_{exc.} = \frac{\hbar^2}{2\mathcal{M}_I} J(J+1)$$



Rotational bands of Hoyle + alpha, ${}^8\text{Be}+{}^8\text{Be}$?

$$E_{exc.} = \frac{\hbar^2}{2\mathcal{M}_I} J(J+1)$$



0_6^+ state : energy gain due to condensation
Rotating alpha dropped at the lowest orbit.

Local condensate \rightarrow complete condensate

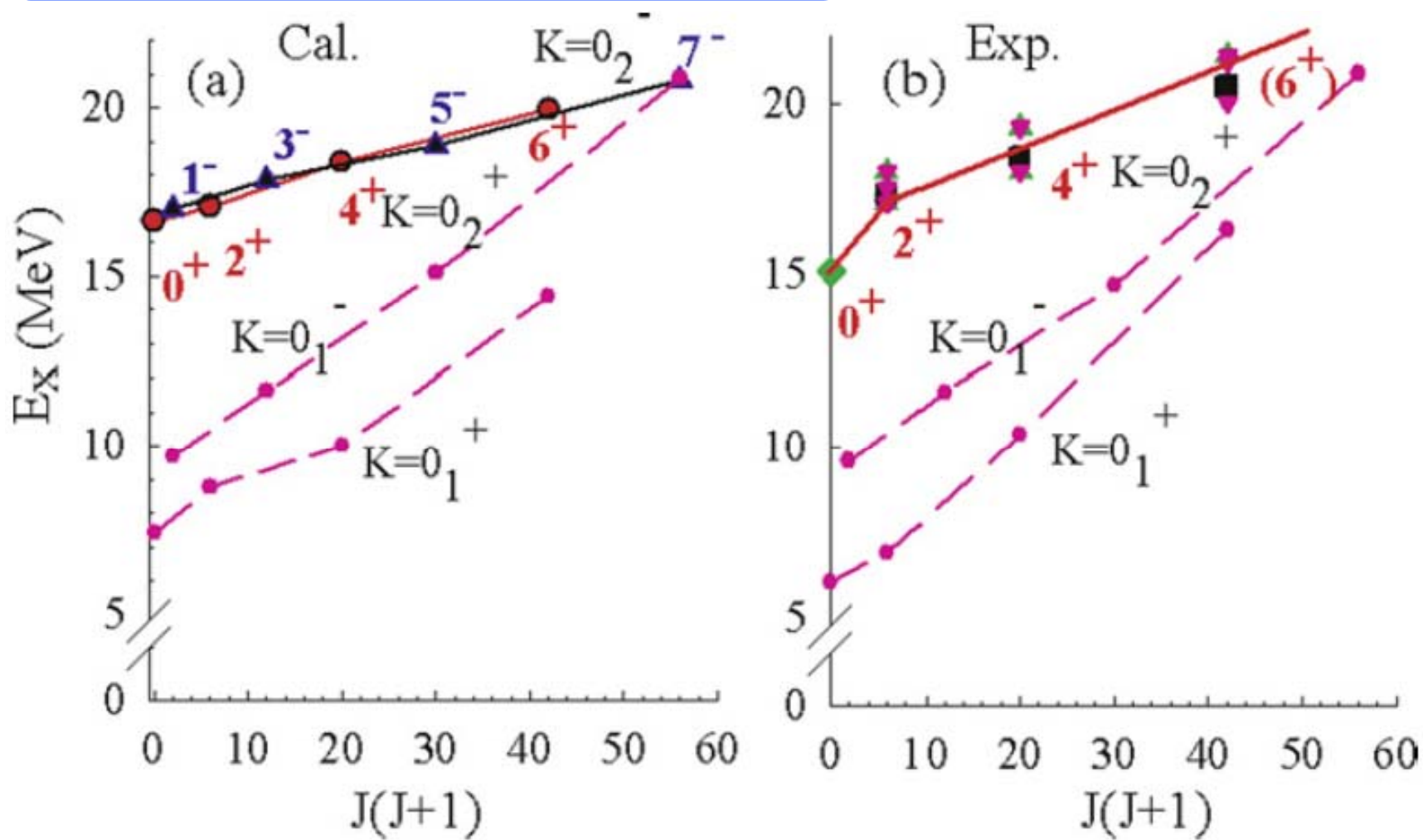
Y.F. S. Ohkubo et al., in preparation.

num.: S^2 -factor of alpha + Hoyle

Momentum inertia is reduced.

Signature of superfluidity ??

Rotational band of Hoyle + alpha



Hoyle + alpha, 2-body scattering solutions.

**Momentum inertia is reduced.
Signature of superfluidity ??**

Summary

Investigation of loosely bound alpha gas states in heavier nuclei than ^{12}C .

- **More α -particle condensate states very likely to exist.**
Analogue state in ^{16}O to the Hoyle state (found with 4α OCM calc.)
as the sixth 0^+ state
Assigned to **15.2 MeV state?**
More experimental information is needed.
- **Hoyle analogs for non-zero spin states are promising.**
likely Hoyle + alpha rotational band
sign of condensate
Problem is continuum mixing

On going issue: beyond bound state approximation

4-alpha CSM (Complex Scaling Method) with T2K-Tsukuba (up to 512cpu's)

Thanks

to my Collaborators

Taiichi Yamada (Kanto Gakuin Univ.)

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Gerd Röpke (Rostock Univ.)

Masaaki Takashina (RCNP)

Tomotsugu Wakasa (Kyushu Univ.)

Wolfram von Oertzen (HMI, Berlin)

Shigeo Ohkubo (Kochi women Univ.)

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