Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

An approach to the construction of EOS Database

> Chikako ISHIZUKA (Tokyo Univ. of Sci.)

### Research Theme

- 1. Liquid-Gas Phase Transition in SN environment
- 2. Hyperon EOS Table for SNe
- 3. Database for Metal Poor Star Observations (SAGA)

Outline of Today's Talk 1. Task at Group A03 2. Current Interests 3. Approach to the construction EOS Database



#### **EOS Information for SNe**

| HOME (EOS List) | Tools | Applicat

Contents Menu

#### Shen EOS and its friends

EOS list EOS Edit Tools Applications/Related Issues

Useful Links

Our Group Our Project

comments on the table e.g.) characteristics of used potential and cal. methods etc.

<u>1998 ver.</u>	
constituents:nucleons, alpha, a heavy nucleu	s
methods:Relativistic Mean Field Theory	
using parameter set TM1+Thomas Fermi calc	
physical constants:	constituents,
m_u=931.49432[MeV],	detailed info. of phys. const.
hbarc=197.32705[MeVfm],	rof EOS table data manual
pi=3.14159265358979323846	rei., EUS ladie aala, manual
ref.)Particle Data Book, PRD 50,3-1(1994) p.1	233
comments:Original	
articles:Nucl. Phys. A, 637 (1998) 435	
manuals: <u>quide for EOS1.pdf</u>	
main table = <u>eos1.tab.qz</u>	
table for zero temperature = <u>eos1.t00.qz</u>	
table for zero proton fraction = <u>eos1.Yp0.qz</u>	
<u>2011 ver.</u>	
constituencts:nucleons, alpha, a heavy nucle	us
methods:Relativistic Mean Field Theory	
using parameter set TM1+Thomas Fermi calc	
physical constants:	
m_u=931.49[4MeV],	

8 EOS tables (row data & codes)

4 Links to the other EOS tables

2 EOS tables coming soon

hbarc=197.327[MeVfm],

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

### Tools for SN simulations

- Add lepton contributions, interpolate EOS tables
- EOS plot function coming soon
- Thomas–Fermi config. of inhomo. phase (S. Yamamuro)
- Transform EOS tables from (Ye, rhoB, T) to (Ye, S, T)

### Future Works

- More EOS Tables (S.Furusawa, H.Togashi)
- Nucleosynthesis/Progenitor models
   (N. Izutani, T. Yoshida, H.Umeda)
- Neutrino interactions (K. Sumiyoshi)

The following groups summarize the useful info. for SN simulations as well <u>http://cococubed.asu.edu/code\_pages/codes.shtml</u> <u>http://www.stellarcollapse.org/</u> <u>http://phys-merger.physik.unibas.ch/~hempel/eos.html</u> <u>http://williamnewton.wordpress.com/links/physics-research/</u>

## Background

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

There are lots of works on nuclear EOS. You can find about 300 papers tiled "EOS" in arXiv in this decade, while we have 15 EOS tables.

Lattimer-Swesty EOS, Shen EOS and their friends are still mainly used in astrophysical studies, nevertheless fine EOSs are NOT always tabulated. What causes this situation? The answer is the cost of tabulating EOS for large grids. We should be free from EOS tables...

Worth constructing EOS DB!?

In addition to the above, exp. data and obs. data related to EOS are now increasing as shown in next slides.

# EOS Problem from NS Obs.

- Recent problems of nuclear EOSs for compact objects "The most of nuclear EOS proposed can not support 2Msum NS with strangess or quarks"
- "MR relation derived from NS obs permit only LS EOS"
   Is that true? Why not?



# Constraints on EOS from nuclear physics



#### An Approach to *Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN* **the Construction of EOS DB**

### The point is...

How do we share these information on EOS, NS obs. data and HIC exp. data? What do we care about when we see these data? e.g.)assumptions, physical conditions, calc.

In EOS DB, we categorize each EOS according to -their theo. background such as framework for "high/medium/low" densities and/or temperatures...(1), -"rel./non-rel."...(2), -off course their framework such as LDM, HF, RHF etc...(3)

## What Database can do is ....

SAGA -Stellar Abundances for Galactic Archeology Database-(URL)<u>http://saga.sci.hokudai.ac.jp/wiki/doku.php?id=start/</u> (UK Mirror)http://www.astro.keele.ac.uk/saga/wiki/doku.php

		Last upda	te of database:	2010-03-16 18:0	01:16				
* not working ** Other options do 1	not work.		Quer	У					
Specify X and Y	search example *	reset							
axis			Graph O	ptions					
Xaxis	Category	Li	(Fe/H)	From :	To :	Include 🛟 data with upper limit			
Yaxis	Category log-e			From :	To :	Include 🛟 data with upper limit			
Criterion +	[X/H] [X/Fe]	Be II		From :	To :	Include 🗘 data with upper limit			
itional criteria	Atmospheric Parameters Photometric Parameters	C	ptional C	riterion	Range				
Object	Binary Parameters Position and Velocity	CH C2	0107-5240"	, "CS", "0107"					
Binarity	Isotopic Ratio	N	n	То					
Magnitude	Band 🗧 : From	NH	mag						
Resolution	< R <	01	01						
		F	graphica	l Criterion					
Author	⊙strict ⊖forward agr	Na Na I Mg	hor A vard agro	ex) "Lastname"	7				
Reference	ALL	Mgi	•						
Publication Year	From To								
			Retrieval (	Options					
Display / Page	10 🗘								
Order by**	Object 🗘								
Output Option	single file 🛟								
Histogram Option	separated files Ran	ige	(ne	cessary for hist	gram)				
	history								

## What Database can do is ....

SAGA -Stellar Abundances for Galactic Archeology Database-(URL)<u>http://saga.sci.hokudai.ac.jp/wiki/doku.php?id=start/</u> (UK Mirror)http://www.astro.keele.ac.uk/saga/wiki/doku.php

	Search Result	SAGA SEAR	сн	RE	sι	JLT		Atmospheric parameters adopted for each object and	
	piot restart reset p t all Results : 420							reference	
	# O Object	Reference	[Fe/H]	Teff	logg	[Fe/H]	log-e(Li)		
	1 🗆 BD-09_4604	C.Charbonnel+, AAP, 442, 961, 2005	-1.38	5660	4	⊙-1.38	• 2.121	Retrieved values and	
	2 🗆 BD-10_155	C.Charbonnel+,AAP, 442, 961, 2005	-2.87	5008	3	●-2.87	● 1.018	aveilable deta Choose one	
111110141111111111	3 🗆 BD-10_388	C.Charbonnel+, AAP, 442, 961, 2005	-2.51	6287	3.85	•-2.51	• 2.257	datum if two or more data	
	4 🛛 BD-12_3709	C.Charbonnel+, AAP, 442, 961, 2005	-1.34	5278	3	●-1.34	• 1.268	ere eveileble	
Choose to plot	5 0 BD.14 5890	C.Charbonnel+, AAP, 442, 961, 2005	-2.07	4885	3	⊙-2.07	●1.025	are available.	
this object	5 55-14_5050	B.W.Carney+, AJ, 125, 293, 2003	-2.01	4840	2.1	<b>○-2.01</b>			
	6 🗆 BD-15_6355	C.Charbonnel+, AAP, 442, 961, 2005	-1.83	6349	4	●-1.83	• 2.351		
	7 BD-17_6692	C.Charbonnel+, AAP, 442, 961, 2005	-1.87	5065	3	•-1.87	• 0.911		
<b>新国新作用和相相</b> 相		M.Spite+,AAP, 430, 655, 2005	-3.06	4750	1.4	○-3.06	⊙ 0.75		
		D.L.Burris+, ApJ, 544, 302, 2000	-3	4575	1.4	0-2.93			
		R.Cayrel+, AAP, 416, 1117, 2004	-3.06	4750	1.4	○-3.06			
		J.A.Johnson+, ApJS, 139, 219, 2002	-2.9	4600	0.95	●-3.03			
		A.McWilliam+,AJ, 109, 2757, 1995				0-2.91			
		J.Melendez+,ApJ, 575, 474, 2002	-2.87	4683	1.7	0-2.87			
		T.V.Mishenina+, AAP, 396, 189, 2002				<b>○-3.01</b>			
	8 O BD 18 5550	T.V.Mishenina+, AAP, 370, 951, 2001	-3.01	4600	0.5	○-3.01			
	6 [] <u>DD-16_5550</u>	M.Spite+,AAP, 455, 291, 2006	-3.06	4750	1.4	○-3.06			
		S.M.Andrievsky+, AAP, 464, 1081, 2007	-3.06	4750	1.4	○-3.06			
T . 1		J.A.Johnson+, ApJ, 658, 1203, 2007	-2.89	4806	1.72	0-2.89			
Links to qui	ck	G.Bihain+,AAP, 423, 777, 2004	-3.01	4668	1.5	0-3.01			
preview files	m	B.W.Carney+, AJ, 125, 293, 2003	-3.11	4820	1.8	<b>○-3.1</b> 1			
HIML form	lat	S.M.Andrievsky+, AAP, 481, 481, 2008	-3.06	4750	1.4	○-3.06			
		J.A.Johnson+, ApJ, 579, 616, 2002				○-3.04			
		J.A.Johnson+, ApJ, 554, 888, 2001	-2.9	4600	0.95	O-3.05			
	9 🔲 BD-20_6718	C.Charbonnel+, AAP, 442, 961, 2005	-1.23	5064	3	⊙-1.23	• 0.953		
RESULTING THE RES	10 D BD 21 2420	R.Smiljanic+, ApJ, 644, L121, 2006	-1.04	5946	3.96	0-1.04	<ul><li>●1.95</li></ul>		
	10 0 00-21_3420	E.Caffau+,AAP, 441, 533, 2005	-1.04	5946	4.41	●-1.04			
	previous         next           1         2         3         4         5         6         7         8         9         10         11         12         13         14         15         16         17         18         19         20           21         22         23         24         25         26         27         28         29         30         31         32         33         34         35         36         37         38         39         40           41         42								

## What Database can do is ....

SAGA -Stellar Abundances for Galactic Archeology Database-(URL)http://saga.sci.hokudai.ac.jp/wiki/doku.php?id=start/ (UK Mirror)http://www.astro.keele.ac.uk/saga/wiki/doku.php



### Expected EOS discussion with EOS DB

Using our classifications, we may find unknown aspects of EOS models on the web as shown in the previous slide, besides we can customize EOS for compact objects within exp./obs. constraints by using EOS DB.

(Lattice QCD, HIC exp. and NS Obs.) +(Bayesian Analysis) = the most probable EOS
The EOS will allow us the following discussions;
1. Internal structure of NS from phenomenological Lagrangian by fitting the above EOS (missing higher order effects or int. strength)
2. NS radius by OV eqn. and observed NS mass An Approach to *Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN* the Construction of EOS DB

Compiled Object

Nucl. Theo. EOS/HIC Exp. SO, L/NS Obs.

Method

Database: MySQL+CSV

Search&Plot web system: Perl/cgi

Compilation by hand

Data

taken from papers (Fig./Tab./sentences) with graph-read soft at present, however in future to be provided from the authors of published papers

> "We need your kind cooperation in constructing our EOS DB"

# Notations for symmetry energy

$$E(\rho, \delta) = E(\rho, 0) + E_{sym}(\rho)\delta^{2} + o(\delta^{4})$$

$$E(\rho, 0) = E(\rho_{0}, 0) + \frac{K_{0}}{2}\varepsilon^{2} + o(\varepsilon^{3})$$

$$E_{sym}(\rho) = E_{sym}(\rho_{0}) + L\varepsilon + \frac{K_{sym}}{2}\varepsilon^{2} + o(\varepsilon^{3})$$

$$K_{0} = 9\rho_{0}^{2}\frac{\partial^{2}E(\rho, 0)}{\partial\rho^{2}}\Big|_{\rho=\rho_{0}}$$

$$\delta = (\rho_{n} - \rho_{p})/\rho$$

$$\varepsilon = (\rho - \rho_{0})/3\rho_{0}$$

$$S_{0} = E_{sym}(\rho_{0})$$

$$L = 3\rho_{0}\frac{\partial E_{sym}(\rho)}{\partial\rho}\Big|_{\rho=\rho_{0}} = (3/\rho_{0})P_{0}$$

$$K_{sym} = 9\rho_{0}^{2}\frac{\partial^{2}E_{sym}(\rho)}{\partial\rho^{2}}\Big|_{\rho=\rho_{0}}$$

$$K_{\tau} \approx K_{sym} - 6L$$

# Task at Group A03 as an "web-master"

### Maintaing Group A03 webpages (Prof. Suzuki)

(aspht1.ph.noda.tus.ac.jp/bridge\_a03/ 🟫 - C 🚼 クォークカ学・原子核構造に基づく爆発的天体現象と元素合成 平成20年度から開始された新学術領域研究(研究領域提案型)「素核宇宙融合による計算科学に基づいた重層的物質構造の解明」の 原子核構造に基づく爆発的天体現象と元素合成」についての情報です。 トップ/お知らせ お知らせ **悪**要 •研究会「素核宇融合による計算機物理学の進展 - ミクロとマクロの架け橋 - 」 🚥 研究成果 新学術領域「素核宇宙融合」とHPCI戦略ブログラム分野5に関連した研究会 資料 「素核宇融合による計算機物理学の進展 - ミクロとマクロの架け橋 - 」が開催されます。 ぜひご参加ください。 研究会など 日時:2011年12月3日午後1時開始~12月5日(月)15:00終了(予定) 開催場所:合歓の郷 http://www.nemunosato.com/index.html リンク 主催:新学術領域研究「素核宇宙融合による計算科学に基づいた重層的物質構造の解明」 HPCI戦略プログラム分野5「物質と宇宙の起源と構造」 メンバーのみ 研究会申込および詳細についてはこちらをご覧ください。(2011年9月12日) 第2回「クォーク力学・原子核構造に基づく爆発的天体現象と元素合成」研究会 研究会ホームページに発表ファイルをあげました。こらちをご覧ください。(2011年6月15日) 昨年に引き続き今年度は東大(本郷) にてこちらの要領で研究会を開催いたします。

### Listing up EOSs (Prof. Sumiyoshi)

ふるってご参加ください。(2010年5月10日)

http://aspht1.ph.noda.tus.ac.jp/bridge\_a03/EOS-index.html

ID: eosdb PW: snmatter

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

🛃 – nu

☆ - C'

## Gathering EOS table Summarising their theoretical background e.g.) Physics constants, method...

/bridge\_a03/EOS-index.html
SN2005E
Photo of a supernova from the
Hubble space telescope.

#### EOS Information for SNe

| HOME | EOS List | Tools | Applications/Related Issues | About Us | Contact Us

#### Contents Menu

#### Introduction

EOS list EOS Edit Tools Applications/Related Issues

Useful Links
Our Group
Our Project

This web-site is managed by research programme A03 of Grant-in-Aid for Scientific Research on Innovative Areas "Elucidation of multirayered materical structure based on computational scienece by the fusion of particle physics, nuclear physics, and astro physics". Our challenge is to understand various problems about the origin of matter sucha as element particle interactions, nuclear stuctures, and compact objects, within the great unified framework. We hope that our challenge would reach new research model not only to explain a heavy element nucleosynthsis mechanism and explosive phenomena but also to understand a complex material structure.

Source: Wikimedia

#### O More information about us

#### News & Update Information





Our consensus is S<sub>0</sub>=31-34 MeV and L=50-110 MeV

Now preparing a summary article on outcomes of NuSYM11.

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

Tools for SN simulations

- Add lepton contributions, interpolate EOS tables
- Figure out EOSs coming soon
- Thomas-Fermi config. of inhomo. phase (S. Yamamuro)
- Transform EOS tables from (Ye, rhoB, T) to (Ye, S, T)

os/EOS-tools.html		☆ - C 🚼 - nuc
EOS Information for	or SNe	SN2005E Photo of a supernova from the Hubble space telescope. Source: Wikimedia Commons/NASA
Contents Menu	Add leptons' contribution to EQS	
EOS list EOS Edit Tools Applications/Related Issues	-add leptons and photon to Shen-s <u>ep</u> <u>another fortran code to add leptor</u> -interpolation	tyle EOS tables Is and photons
Useful Links		
<u>Our Group</u> <u>Our Project</u>		

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN



Copyright (C) 2011 Bridge-A03 All Rights Reserved. design by tempnate This page is maintained by Chikako Ishizuka. Last updated at 22nd. Sep., 2011

## What SAGA can do is ....

Workshop "Progress in computational physics---Bridge between MACRO and MICRO 3rd - 5th, Dec, 2011 Ise, Shima, JAPAN

SAGA -Stellar Abundances for Galactic Archeology Database-(URL)<u>http://saga.sci.hokudai.ac.jp/wiki/doku.php?id=start/</u> (UK Mirror)http://www.astro.keele.ac.uk/saga/wiki/doku.php

